

# Stability analysis of the Labrador Current

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# Outline

## 1) Motivation

-> Labrador Current

-> seasonality of EKE along the Labrador Current

## 2) FLAME model

## 3) Linear stability analysis

## 3) Results

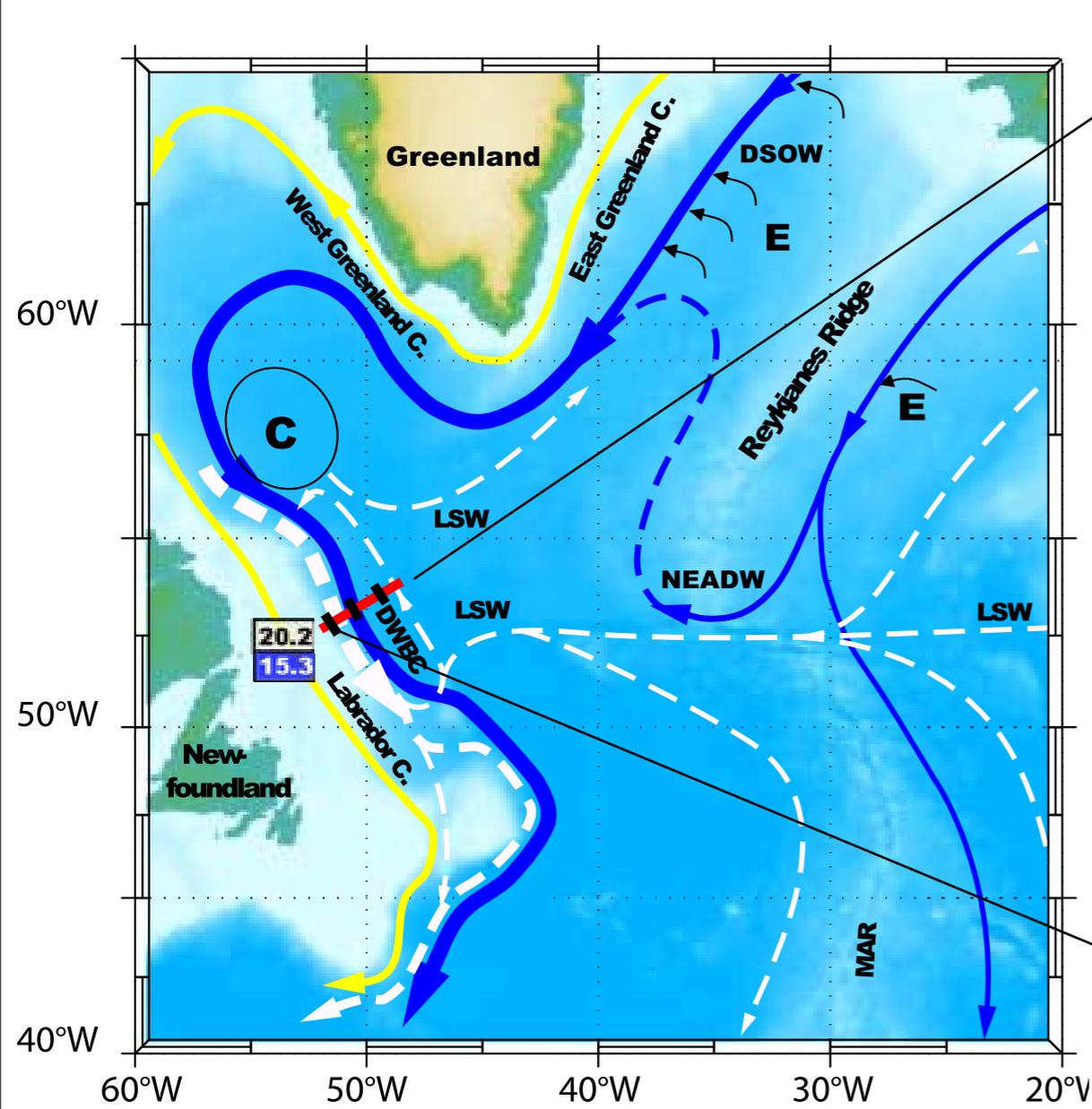
-> interior mode

-> oceanic background conditions

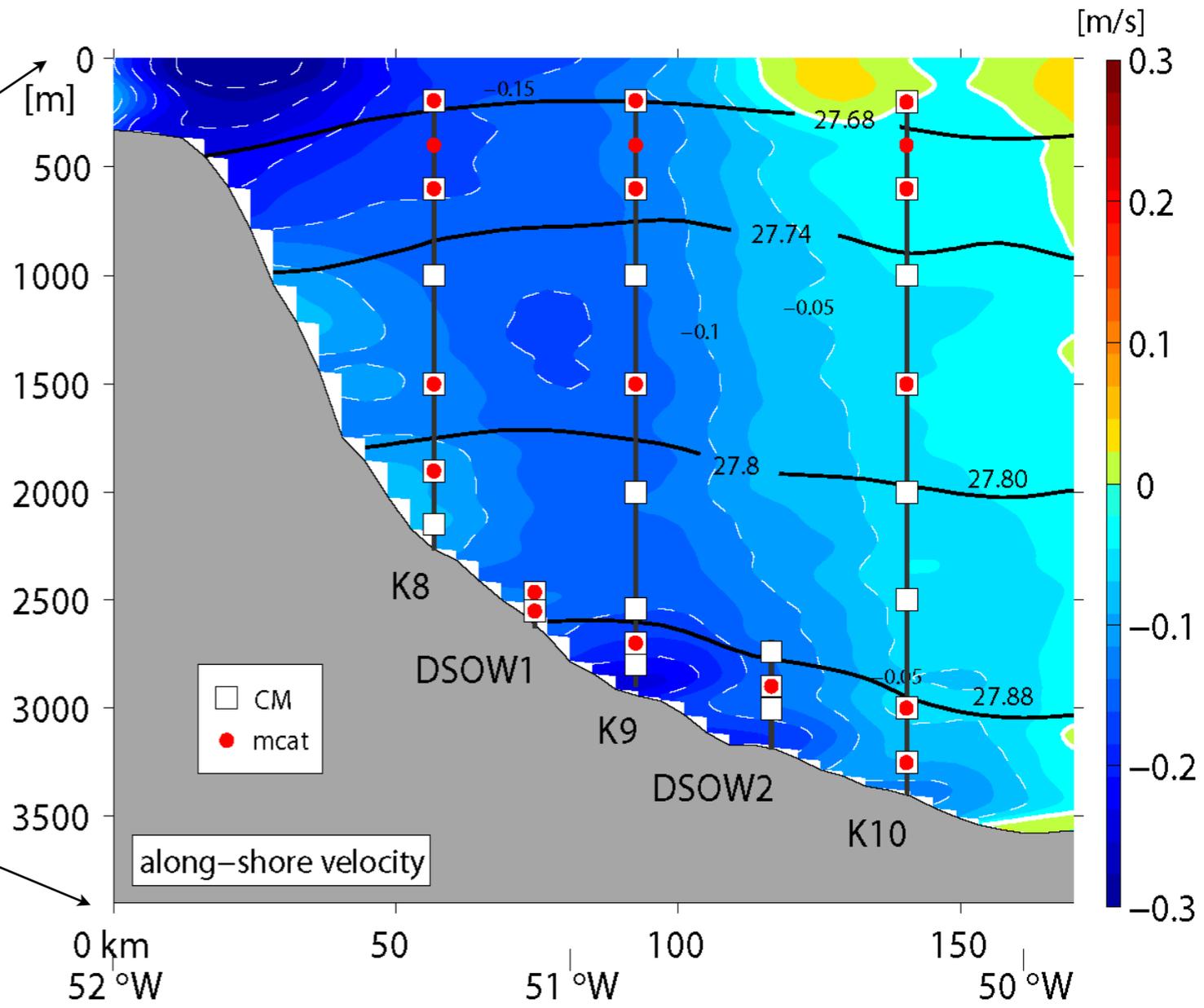
## 4) Discussion

## 5) Summary and conclusion

# Labrador Current



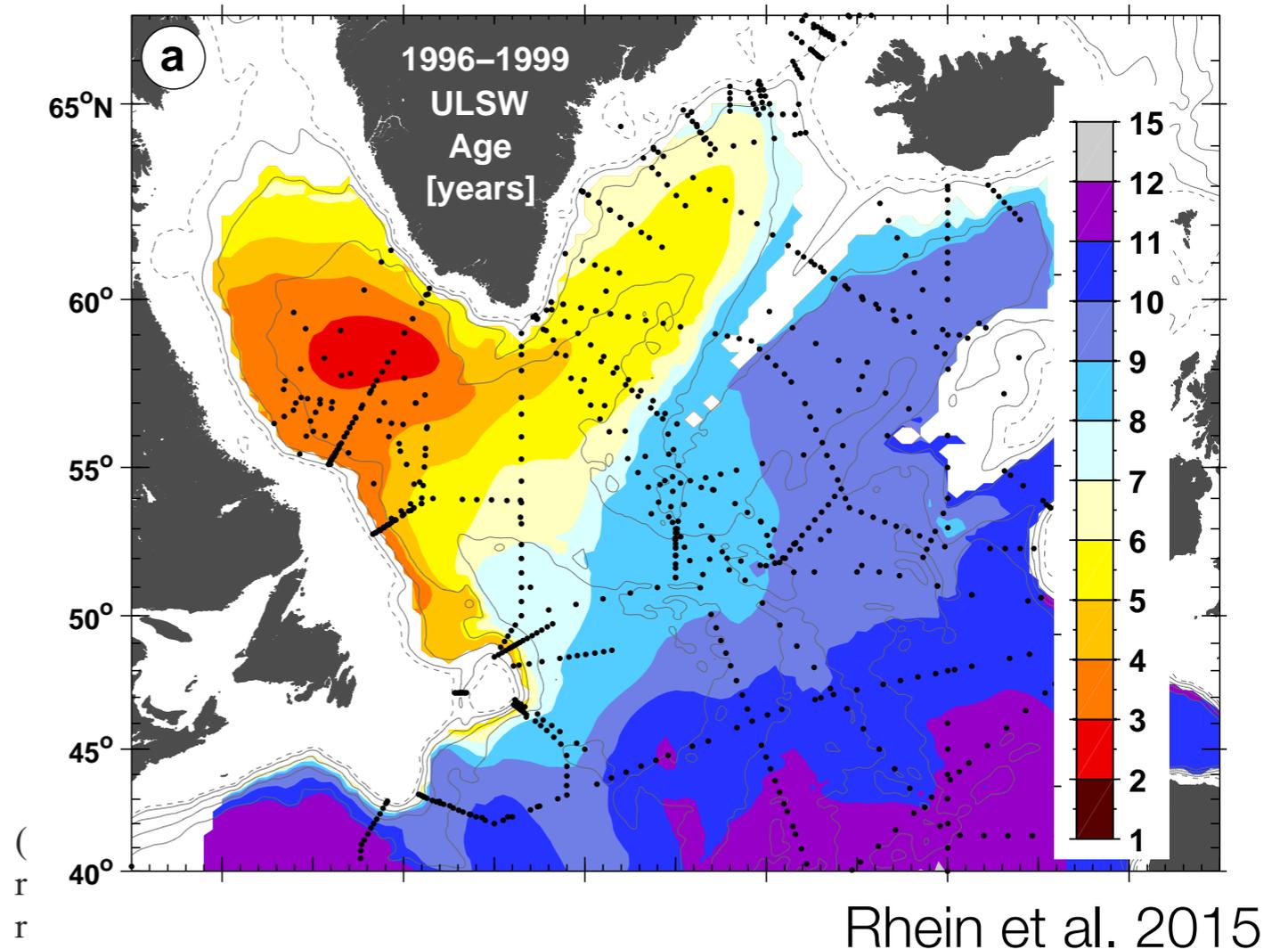
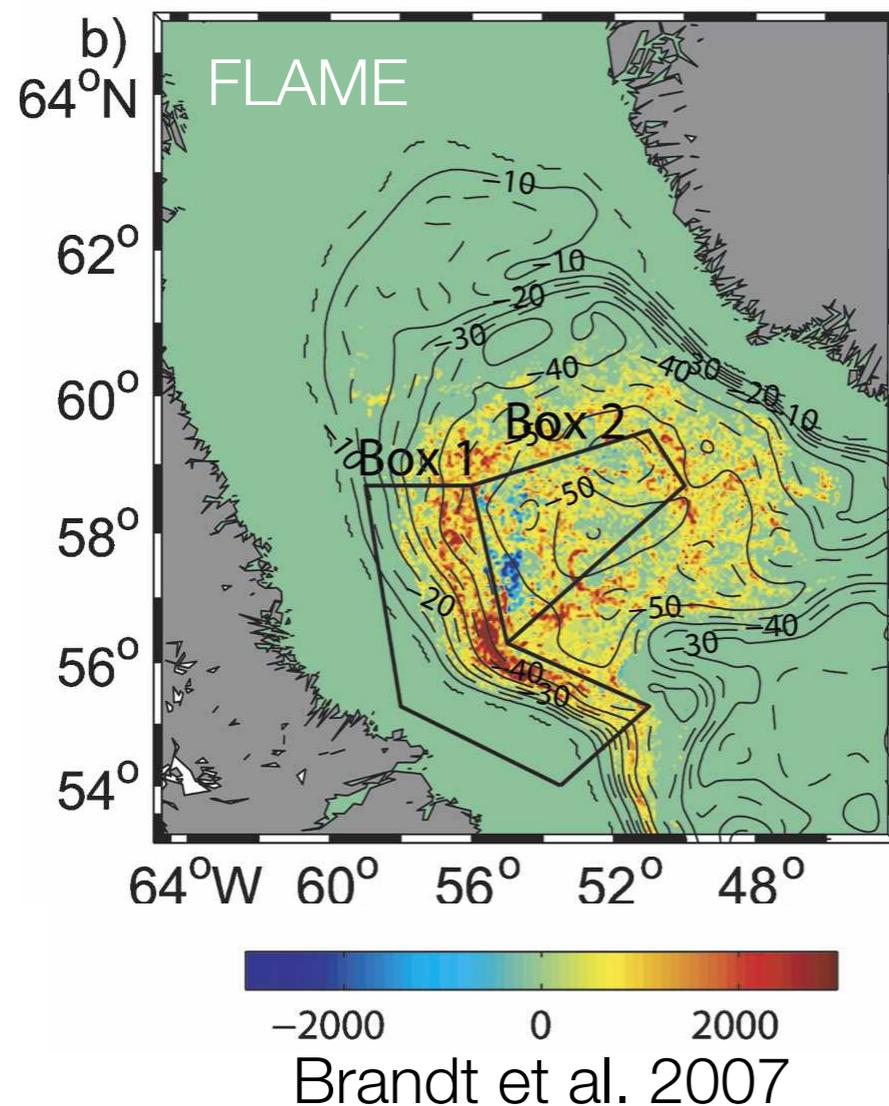
Fischer et al. 2010



Fischer et al. 2014

# LSW transformation and export

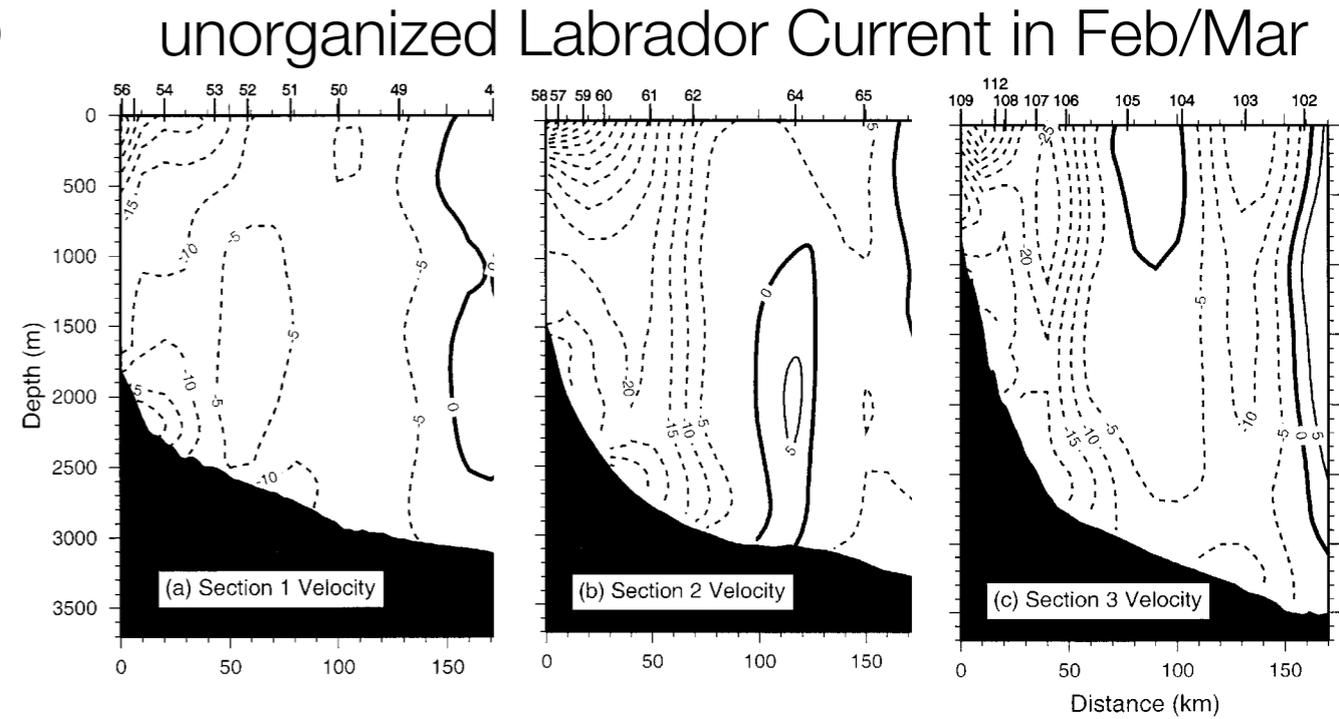
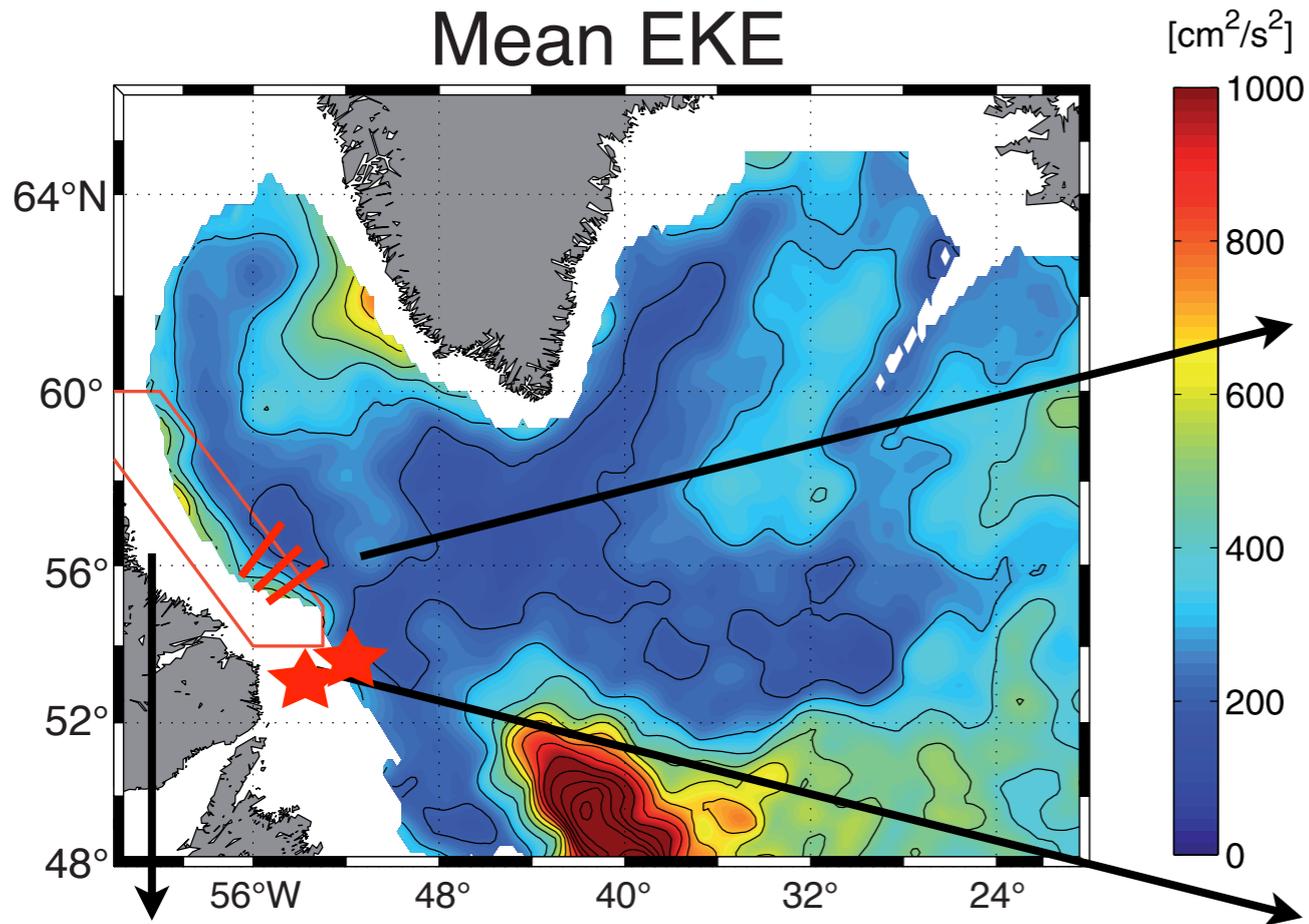
Jan/Mar LSW thickness change



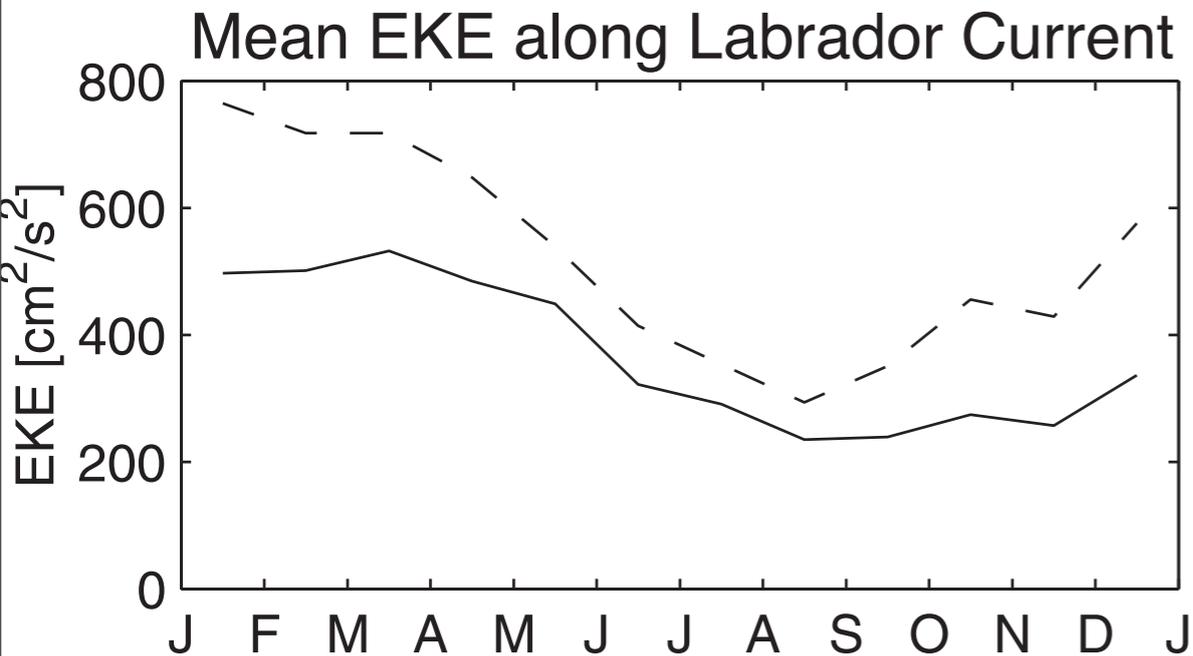
-> Labrador Current is important water mass transformation area  
(Pickart et al. 1997, 2002, Spall and Pickart 2001, Straneo et al. 2002, Cuny et al. 2005, Pickart and Spall 2007, Brandt et al. 2007, Spall 2010)

-> rapid export of ULSW into subtropical gyre  
(Schott et al. 2004, Brandt et al. 2007, Rhein et al. 2015)

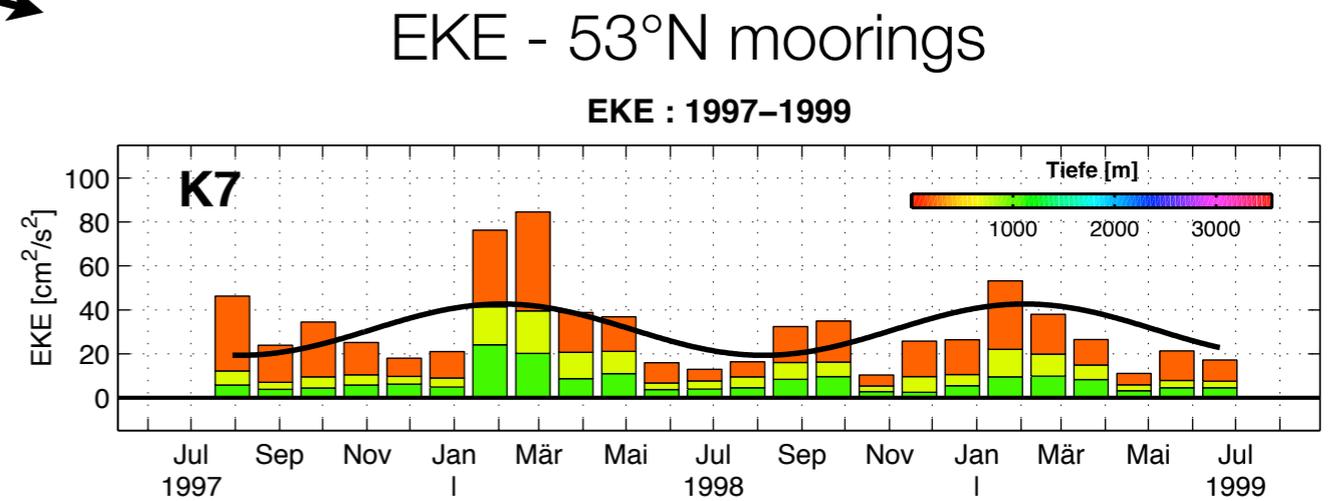
# Seasonality of EKE along LC



Pickart et al. 2002



Brandt et al. 2004

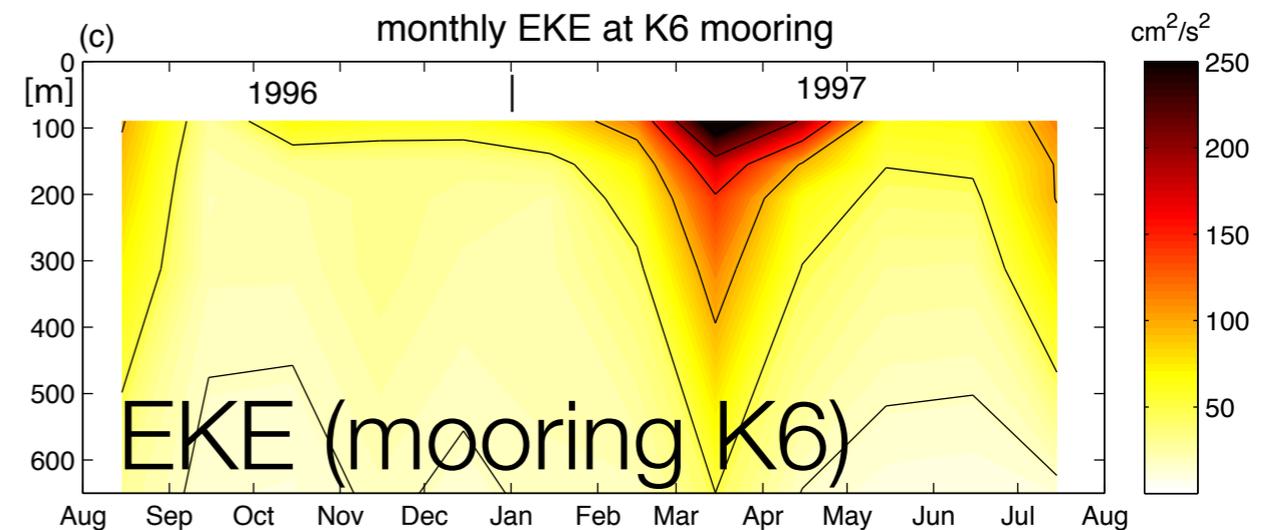
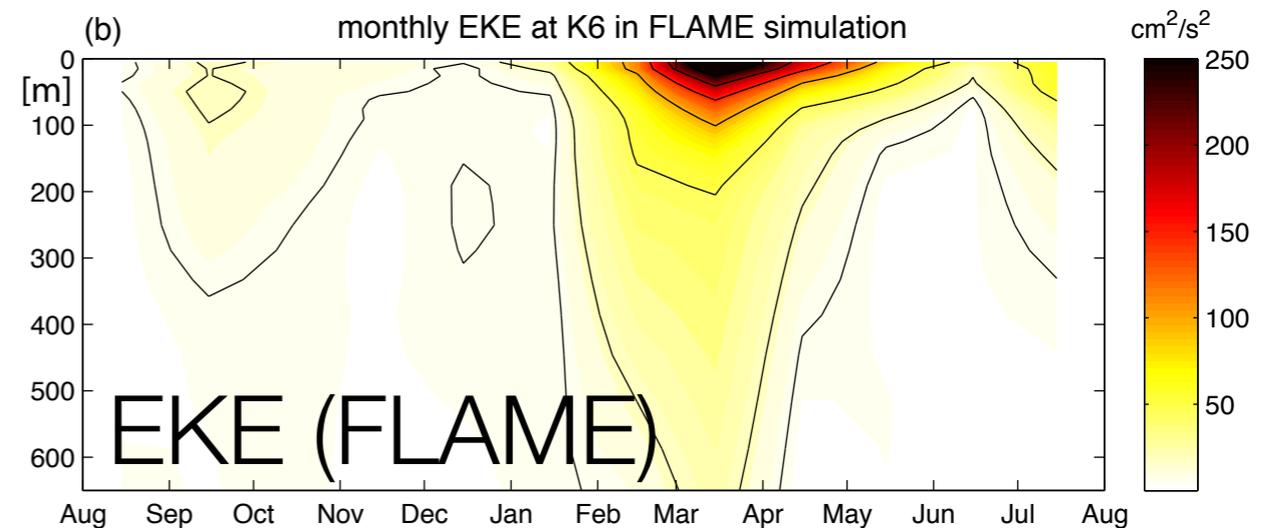


Morsdorf 2001 (Dipl. Thesis)

# Research questions

1) Why does the **Labrador Current** become unstable in **late winter** and what kind of **instability** is at work?

2) How do the **oceanic background conditions** contribute?



# Data / Model / Method

- **moored current data** in Labrador Current (K6, K7, K8 see Fischer et al. 2004, Cuny et al. 2005)
- **1/12°** north Atlantic ocean model from the Family of Linked Atlantic Model Experiments (**FLAME**)
- non-geostrophic **linear stability analysis** to investigate the stability problem

# FLAME Model

- $1/12^\circ \rightarrow dx = 4 - 5 \text{ km}$
- 45 non-equidistant vertical levels
- primitive equations, hydrostatic
- forcing: clim. monthly mean wind and surface fluxes
  
- Czeschel (PhD 2005) improved hydrography and deep convection
  
- > realistic mixed layer depths
  
- Brandt et al. (2007) studied transformation and export of Labrador Sea Water

mixed layer depth March

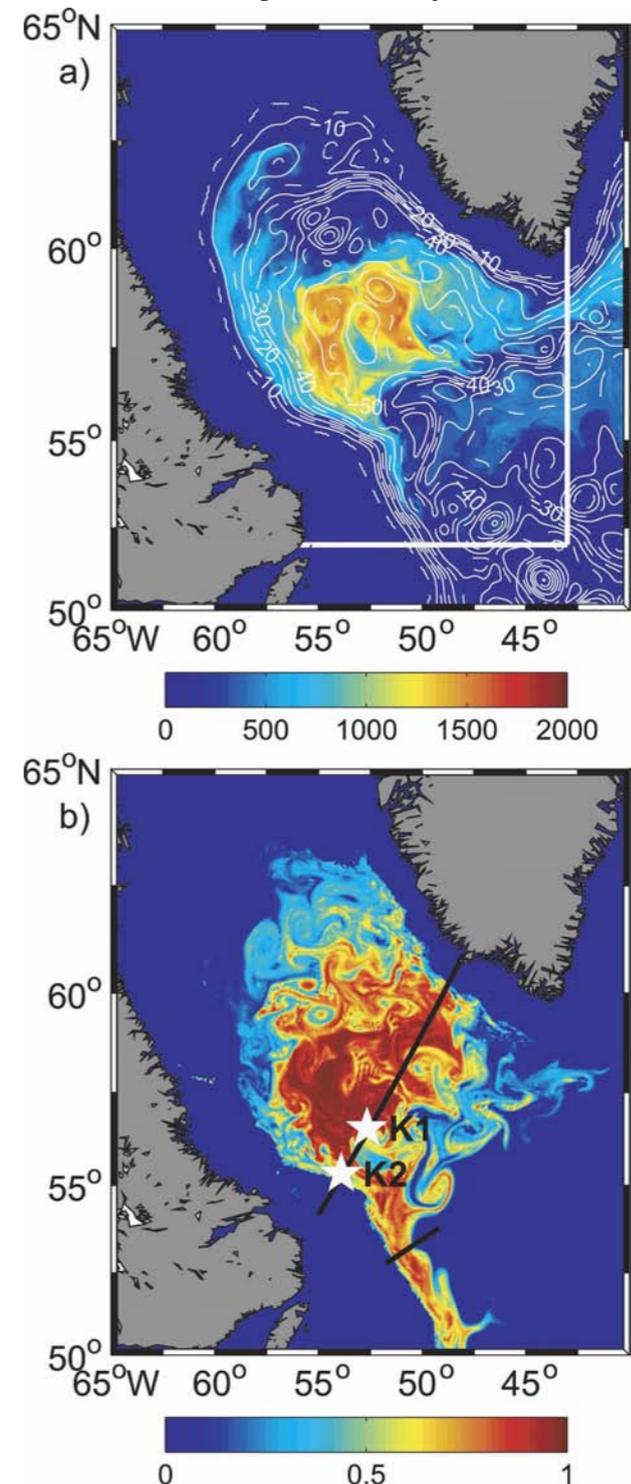
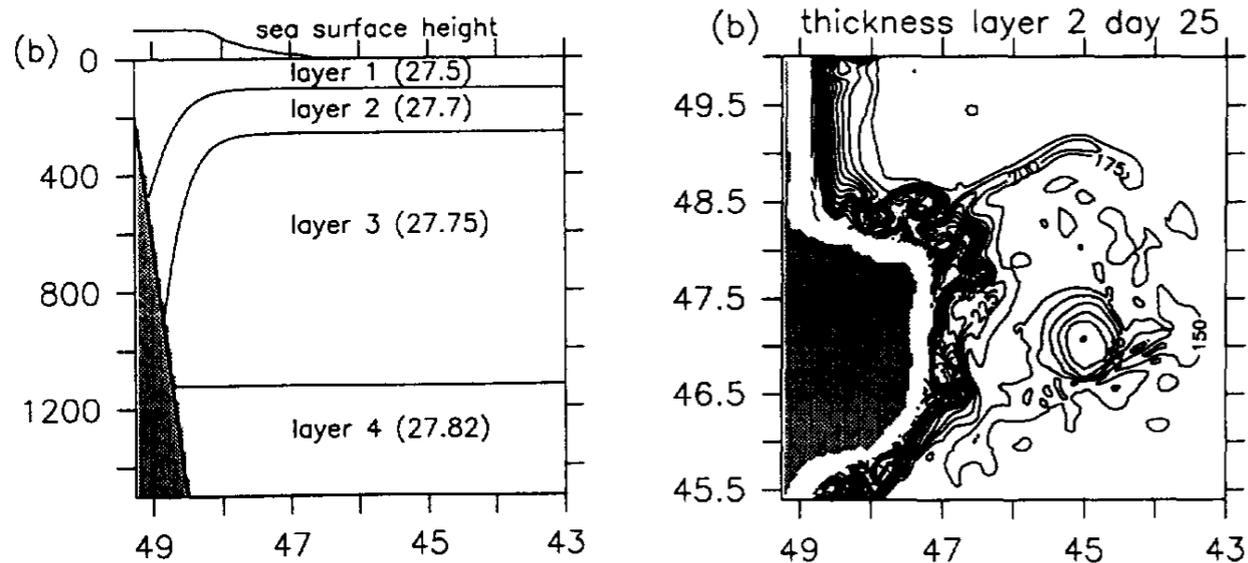


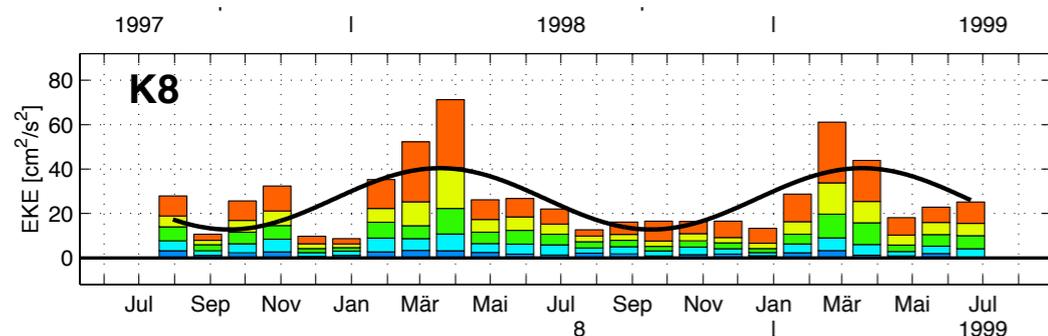
FIG. 1. (a) Simulated mixed layer depth (m; color shading) superimposed on barotropic streamfunction ( $Sv$ ; isolines) during March and (b) concentration of the idealized ventilation tracer in the  $\sigma_\theta$  range 27.77–27.80  $\text{kg m}^{-3}$  during May; (a) also shows the chosen limits for the Labrador Sea region, i.e., the 52°N and the 43°W section, and (b) shows the LC section at about 53°N and the WOCE AR7W section running across the deep Labrador Sea, as well as the locations of moorings K1 and K2 (white stars).

Brandt et al. 2007

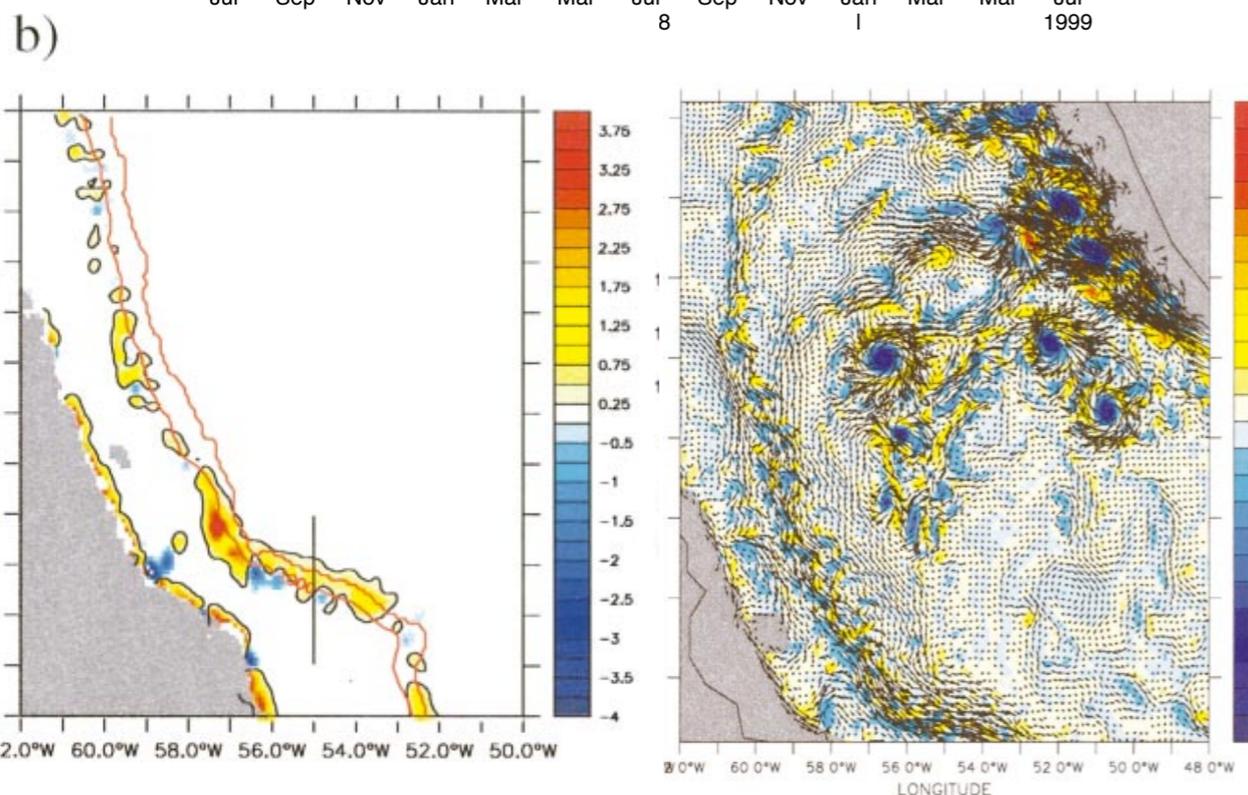
# Unstable Labrador Current



- Pickart et al. 1997 studied uLSW eddy formation process in four-layer primitive equation model



- EKE due to high frequency wind forcing? (Morsdorf, 2001)



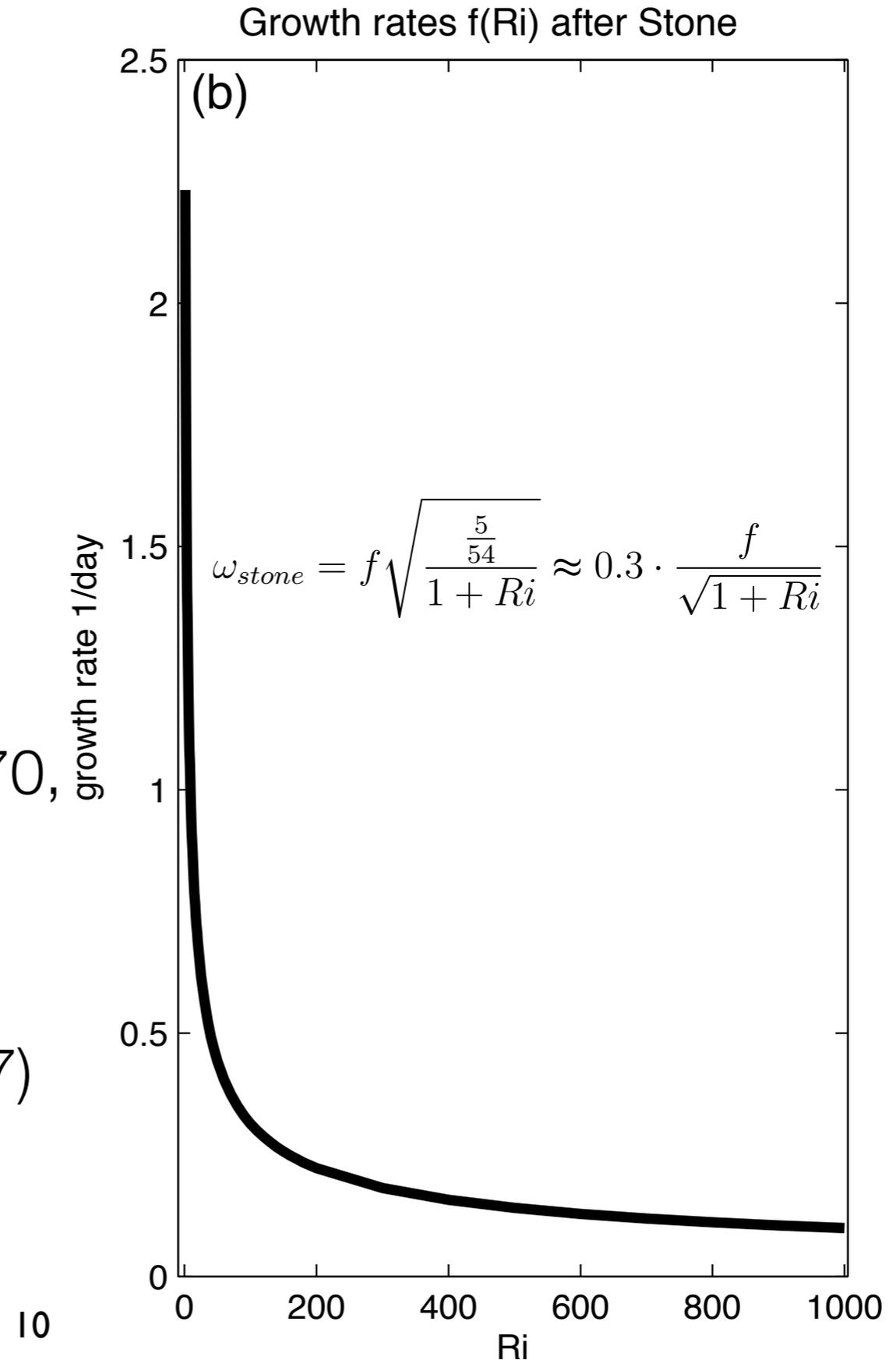
- monthly mean winds to exclude high-frequency winds as a direct source of EKE (Eden and Böning, 2002)

- baroclinic >> barotropic transfer terms (Eden and Böning, 2002)

**-> focus on baroclinic instability problem**

# Baroclinic instability

- growth rate of baroclinic waves
- larger growth rates -> more unstable waves
- growth rate depends on Richardson number (Eady 1949)
- non-geostrophic baroclinic instability problem (Stone 1966, 1970, 1971, 1972)
- growth rate correlates with EKE (Stammer, 1997, Chelton et al. 2007)



# Non-geostrophic linear stability analysis

Set of equations:

$$\partial_t u + \mathbf{u} \cdot \nabla u - fv + f_h w = -\partial_x p \quad \text{momentum}$$

$$\partial_t v + \mathbf{u} \cdot \nabla v + fu = -\partial_y p$$

$$\partial_t w + \mathbf{u} \cdot \nabla w - f_h u = -\partial_z p + b$$

$$\partial_t b + \mathbf{u} \cdot \nabla b = 0 \quad \text{buoyancy}$$

$$c_s^{-2} \partial_t p + \nabla \cdot \mathbf{u} = 0 \quad \text{continuity}$$

- similar to non-geostrophic stability analysis of Stone (1971)
- slightly more general, more details in Thomsen et al. 2014 or Brüggemann and Eden 2014

# Linear stability analysis

Linearize:

$$\mathbf{u} = \mathbf{U} + \mathbf{u}', \quad b = B + b', \quad p = P + p' \quad \text{with } \mathbf{U} = (U_1(z), U_2(z), 0)^T$$

basic state + first order perturbation equation

-> look for wave solutions:

$$u = u_0(z) \exp i(\omega t - kx - ly)$$

- eigen system, discretize and solve it numerically for given wavenumbers  $k$  and  $l$ :

$$\omega \mathbf{a}(z) = \mathbf{M}(z) \mathbf{a}(z) \quad \mathbf{a}(z) = (u_0, v_0, w_0, b_0, p_0)^T$$

# Linear stability analysis

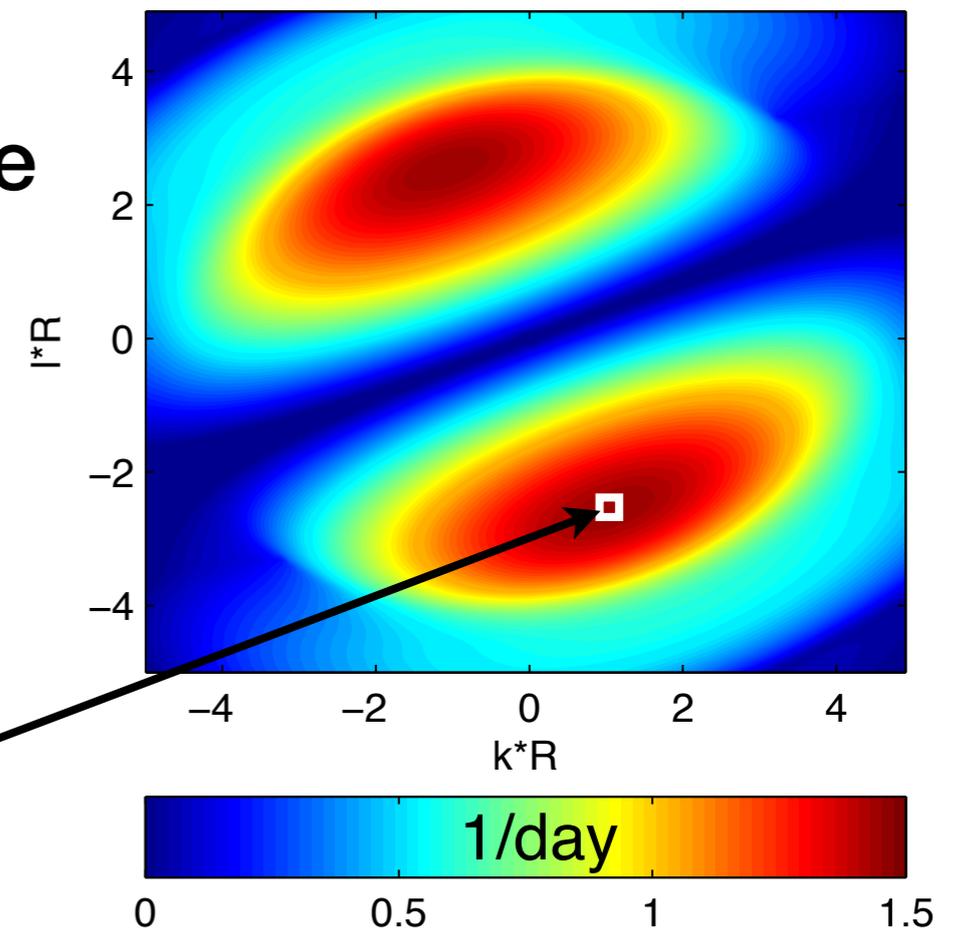
eigen system, amplitude vector  $\mathbf{a}(z)$ , eigenvalue  $\omega$

$$\omega \mathbf{a}(z) = \mathbf{M}(z) \mathbf{a}(z)$$

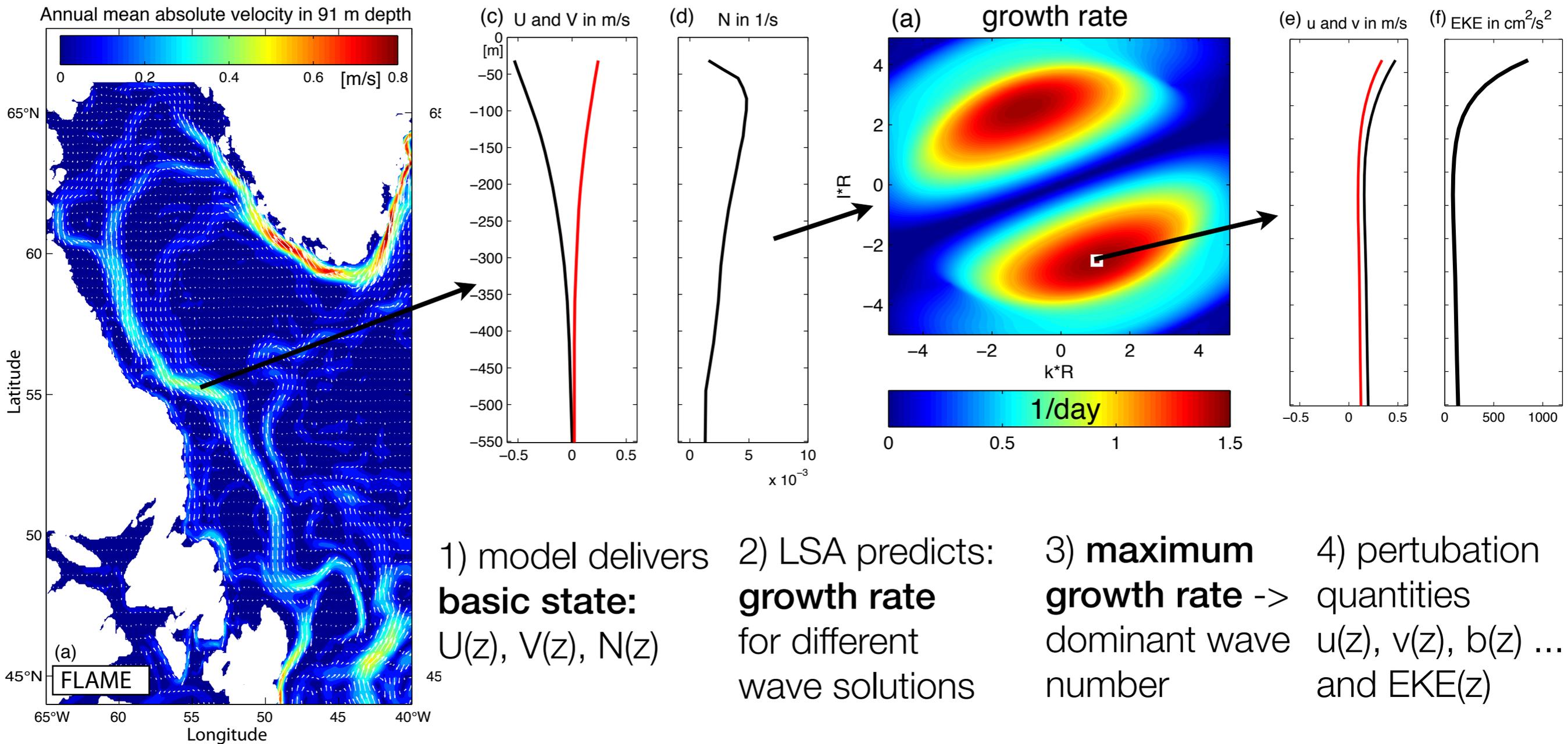
$$\mathbf{a}(z) = (u_0, v_0, w_0, b_0, p_0)^T$$

- imaginary part of eigenvalue -> **growth rate**
- real part of eigenvalue -> **phase velocity**
- positive growth rate -> growing waves
- **pick fastest growing wave solution**

(a) growth rate

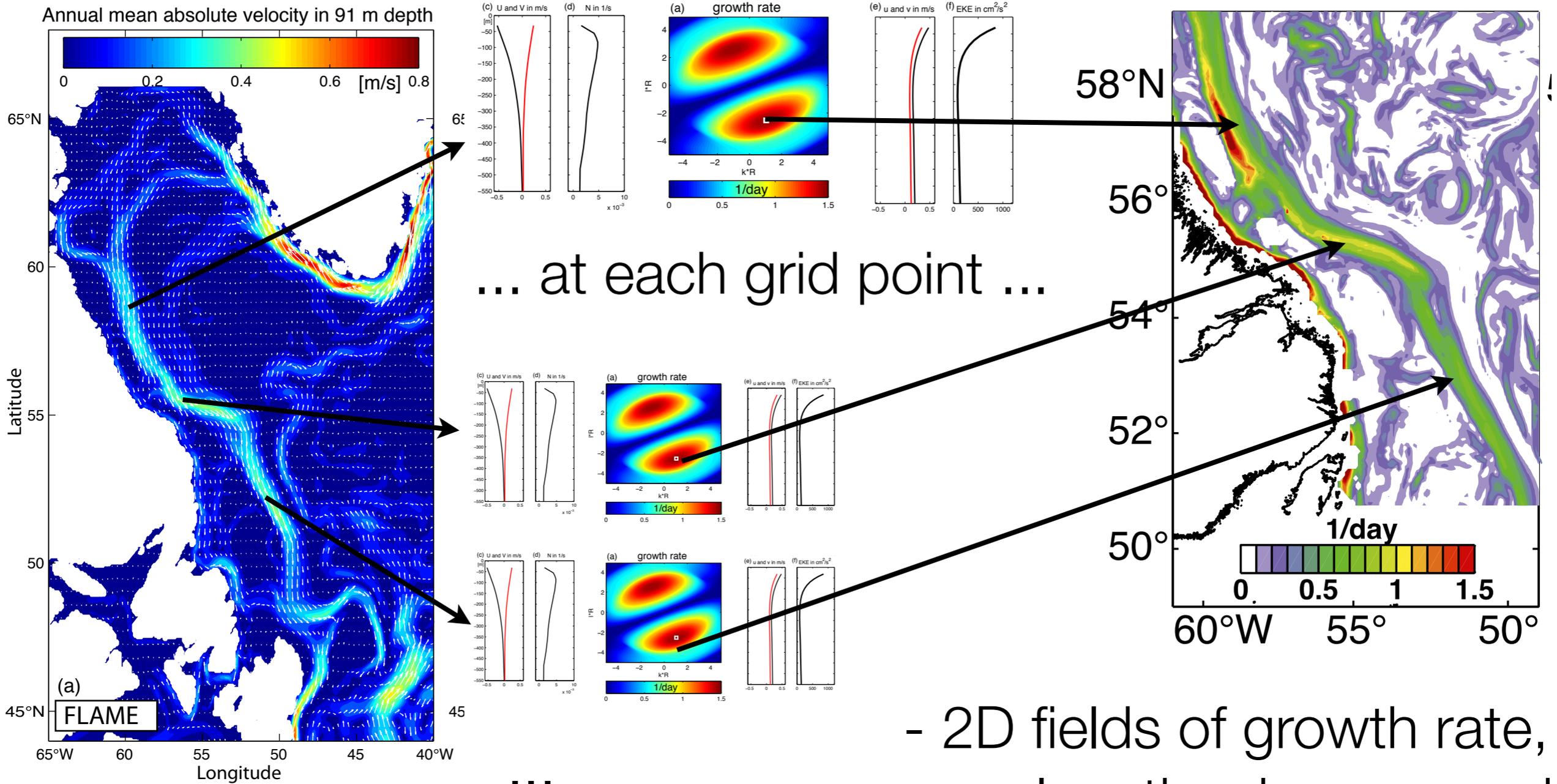


# Strategy: Example profile



- EKE scaling according to Killworth (1997), Eden (2011, 2012), Vollmer and Eden (2013)

# Mapping



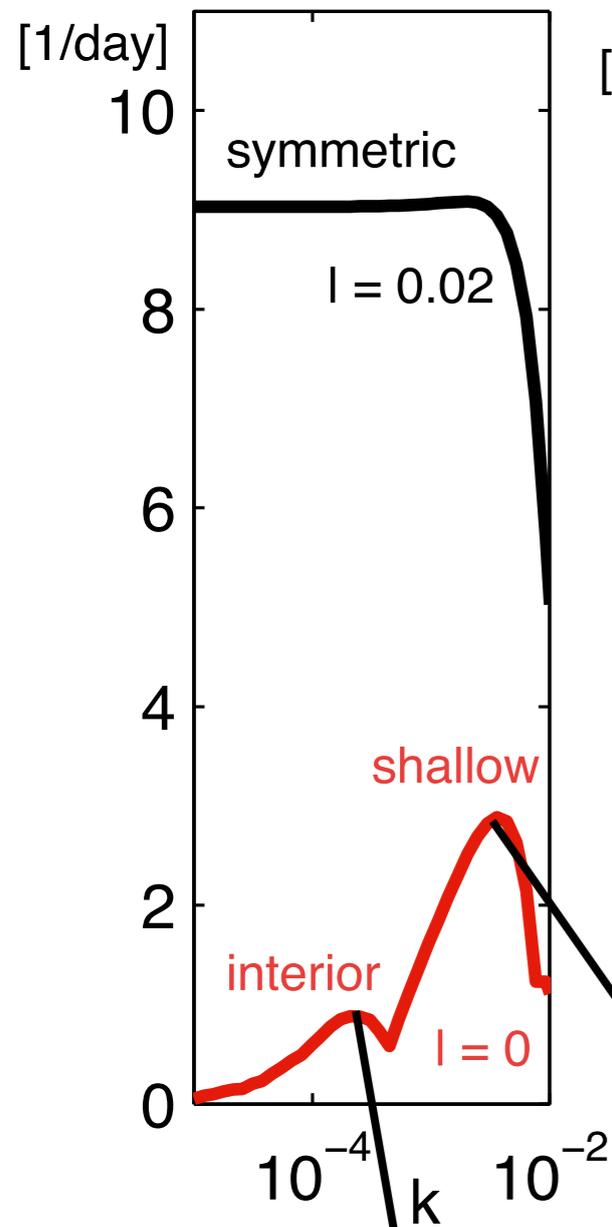
- 2D fields of growth rate, wavelength, phase speed,
- 3D fields of  $u'$ ,  $v'$ , EKE, eddy fluxes, ...

# Results

Linear stability analysis

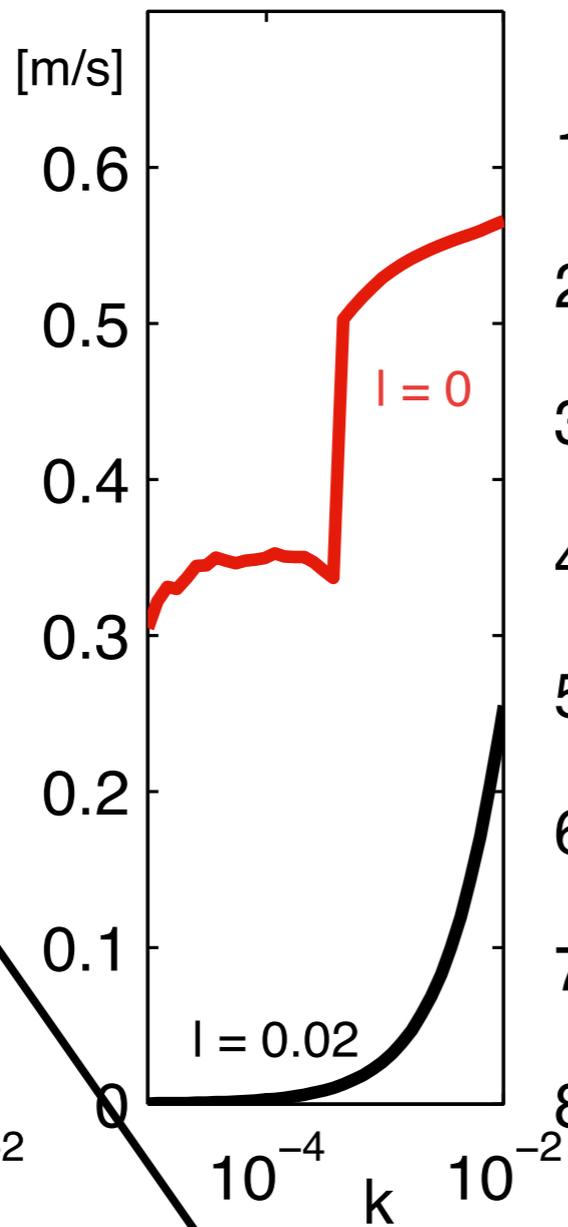
# Interior, shallow and symmetric mode

(a) growth rate



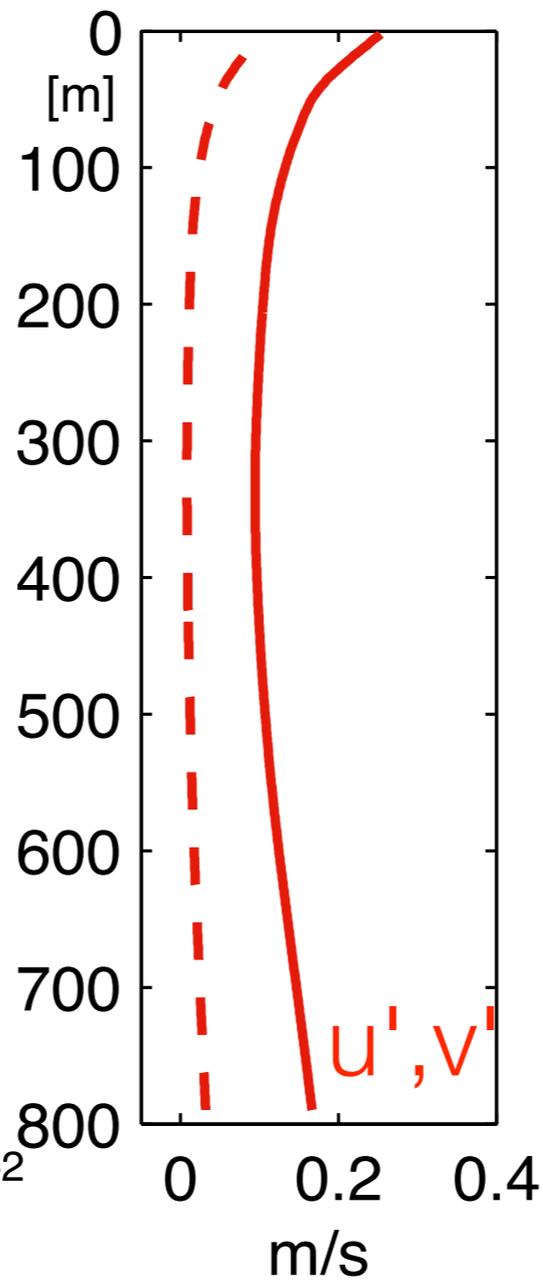
O(30-45 km)

(b) phase velocity

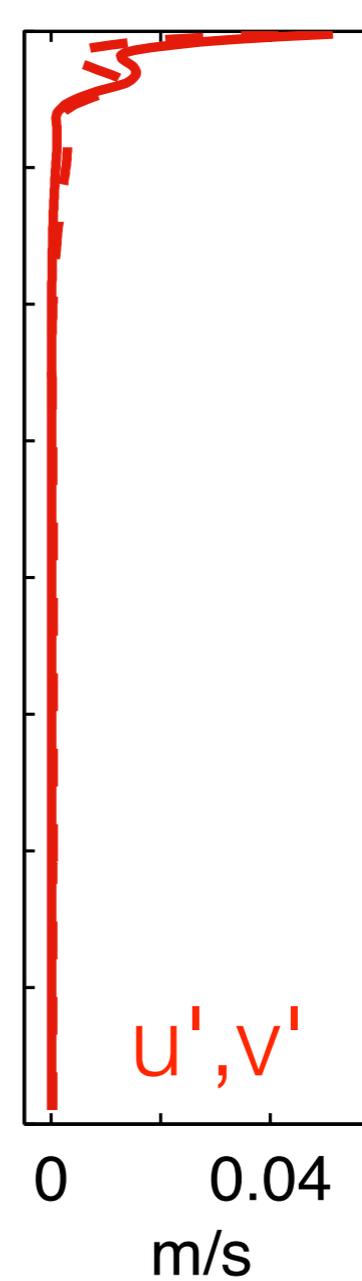


O(1 km)

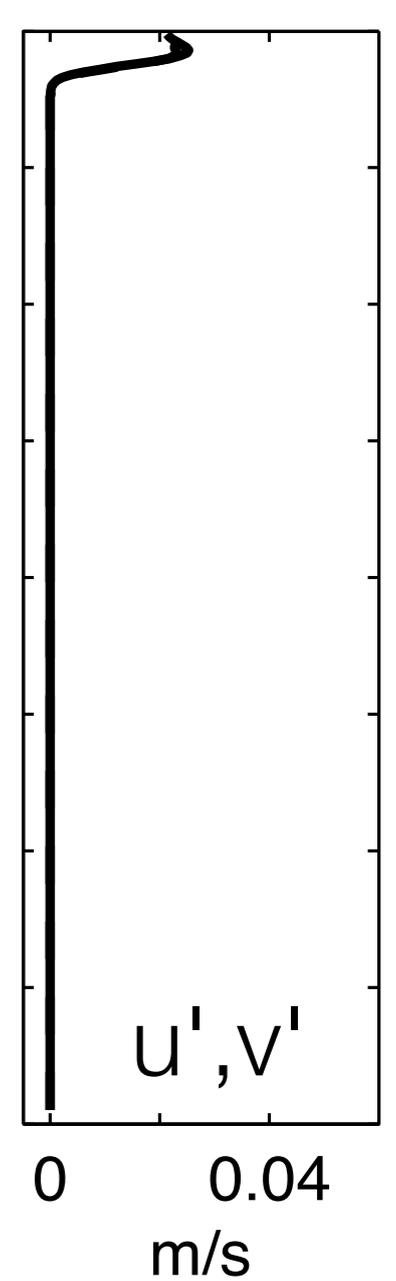
(c) interior



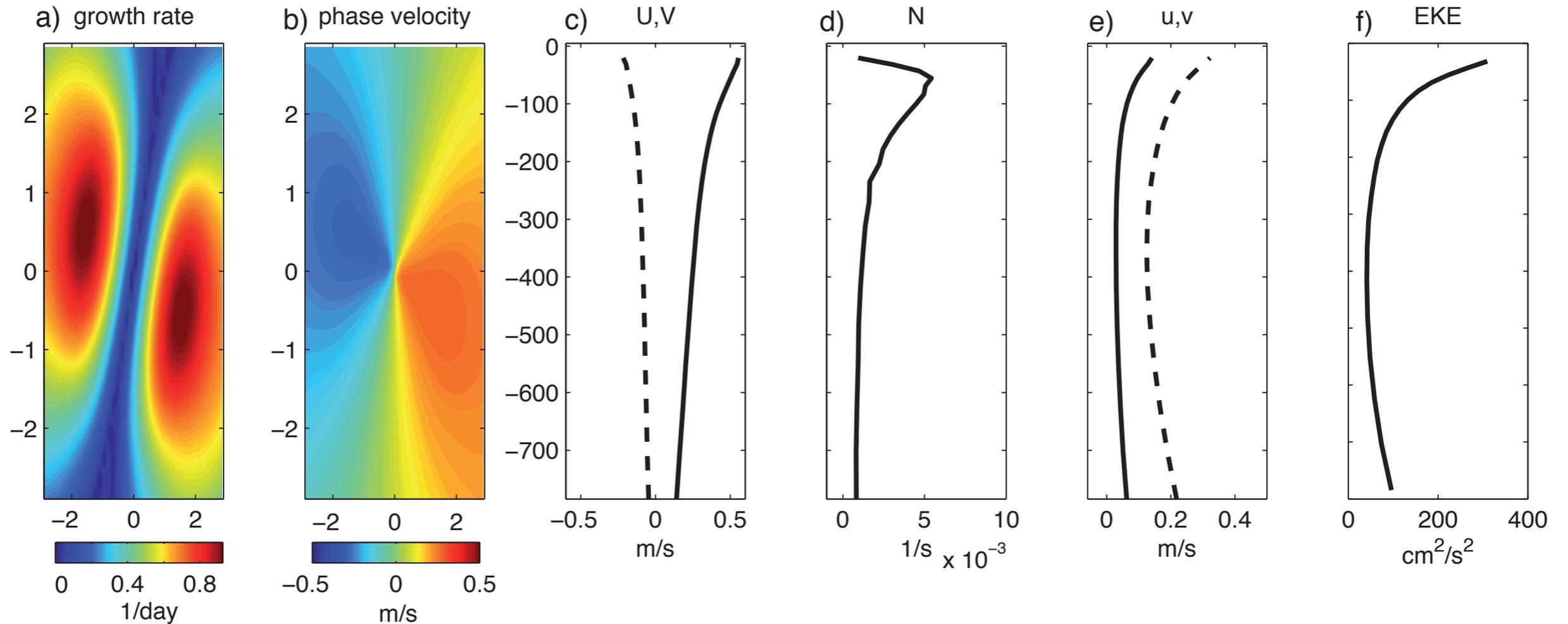
(d) shallow



(e) symmetric



# Interior mode



- wavelength  $\sim$  30 - 45 km as in model
- surface intensified EKE structure as in observations and model
- > responsible for enhanced EKE in model simulations

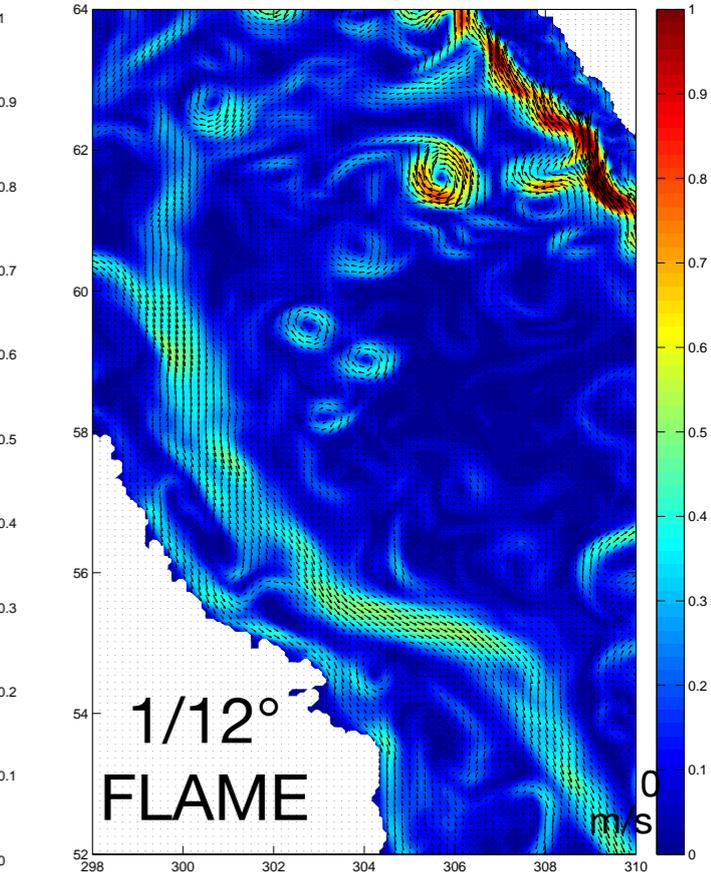
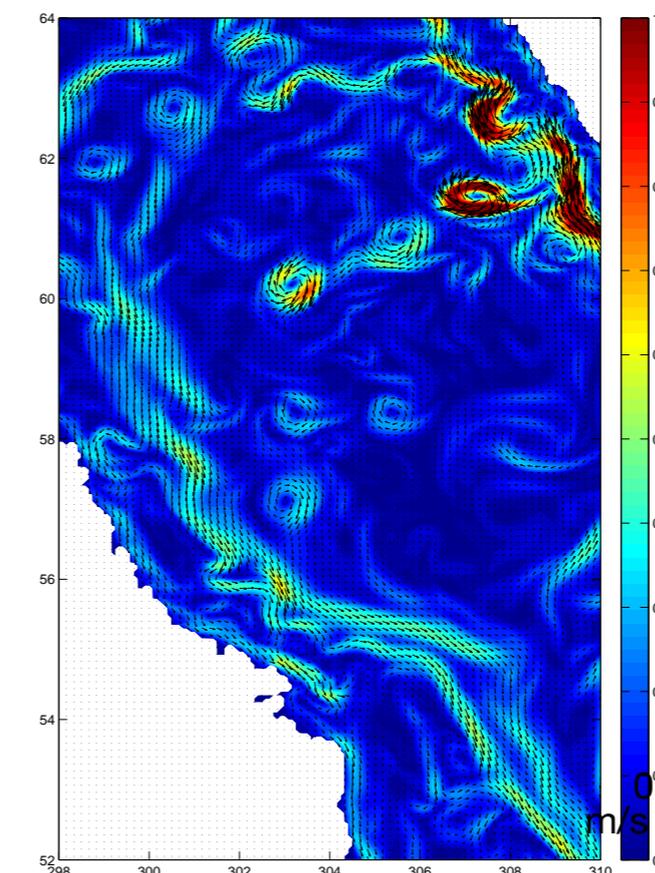
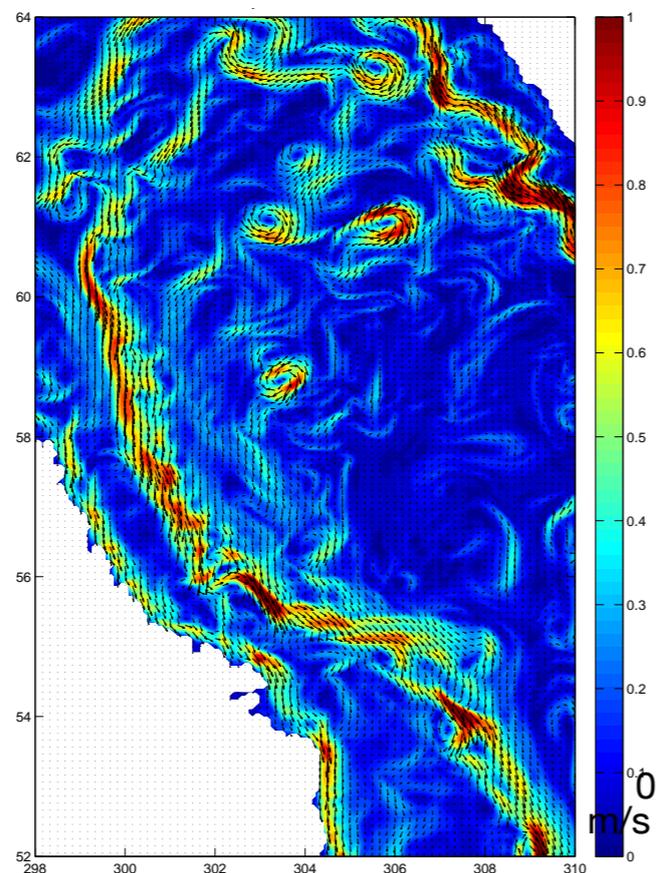
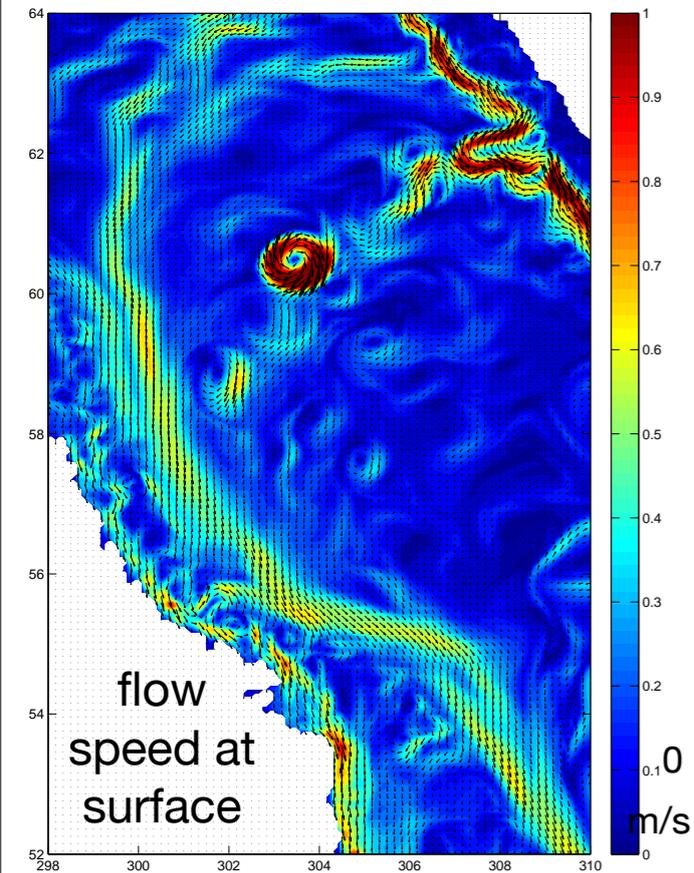
# Unstable Labrador Current in March

December 15

March 15

July 15

August 15



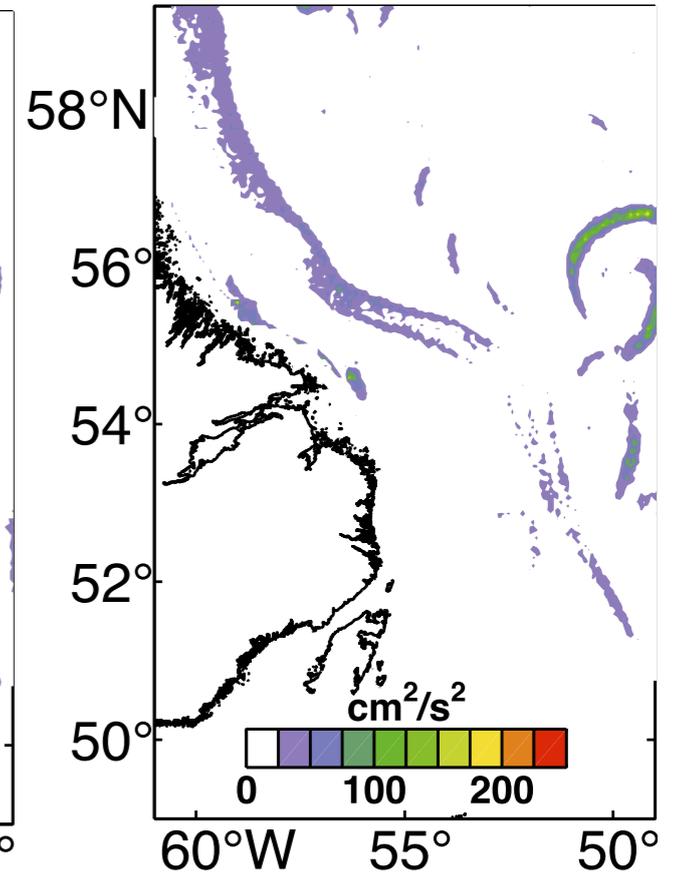
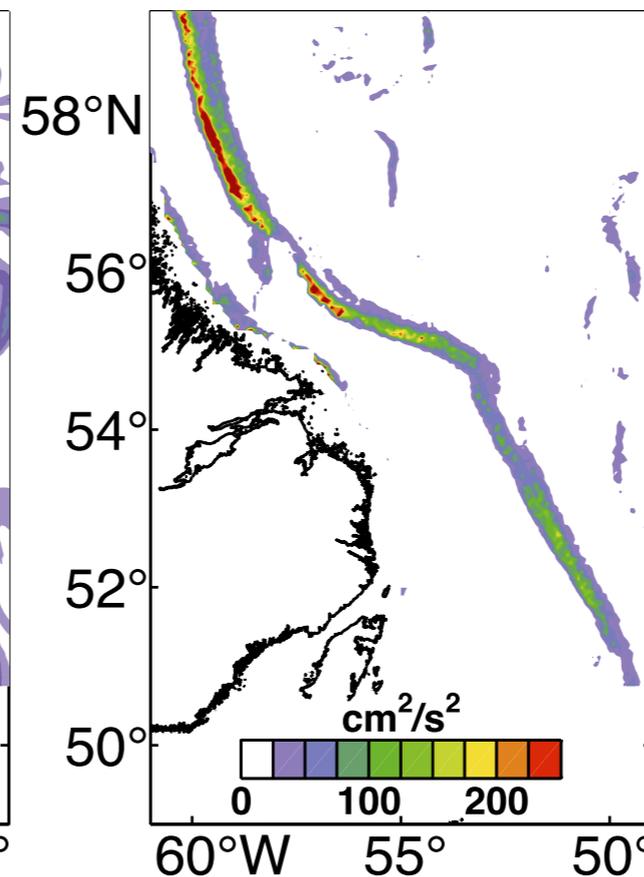
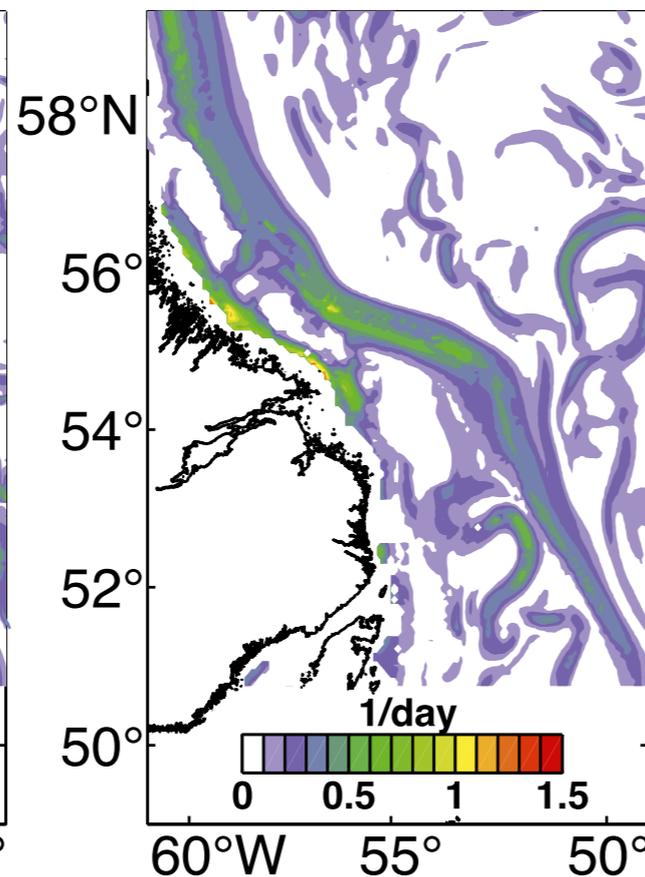
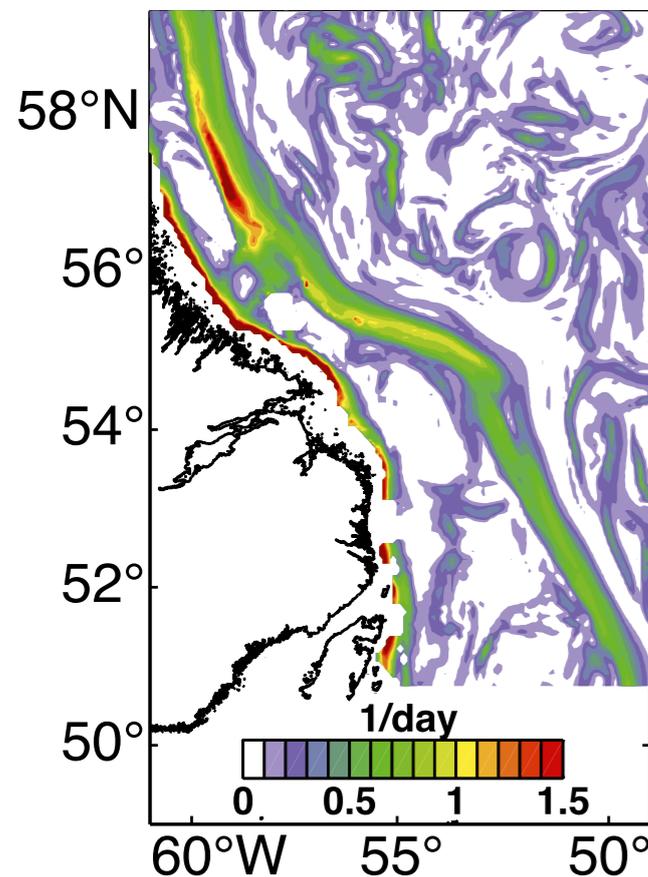
# Predicted growth rates and EKE

(a) growth rate in March

(b) growth rate in Sept.

(c) EKE in March

(d) EKE in Sept.



- > 3 times higher growth rates of interior mode in March compared to September
- > enhanced local EKE production in Labrador Current during March

# Background conditions

- weak stratification
- enhanced vertical shear

-> small Richardson numbers ( $Ri < 5$ ) in late winter

-> growth rates of baroclinic waves increase (Eady 1949, Stone 1970)

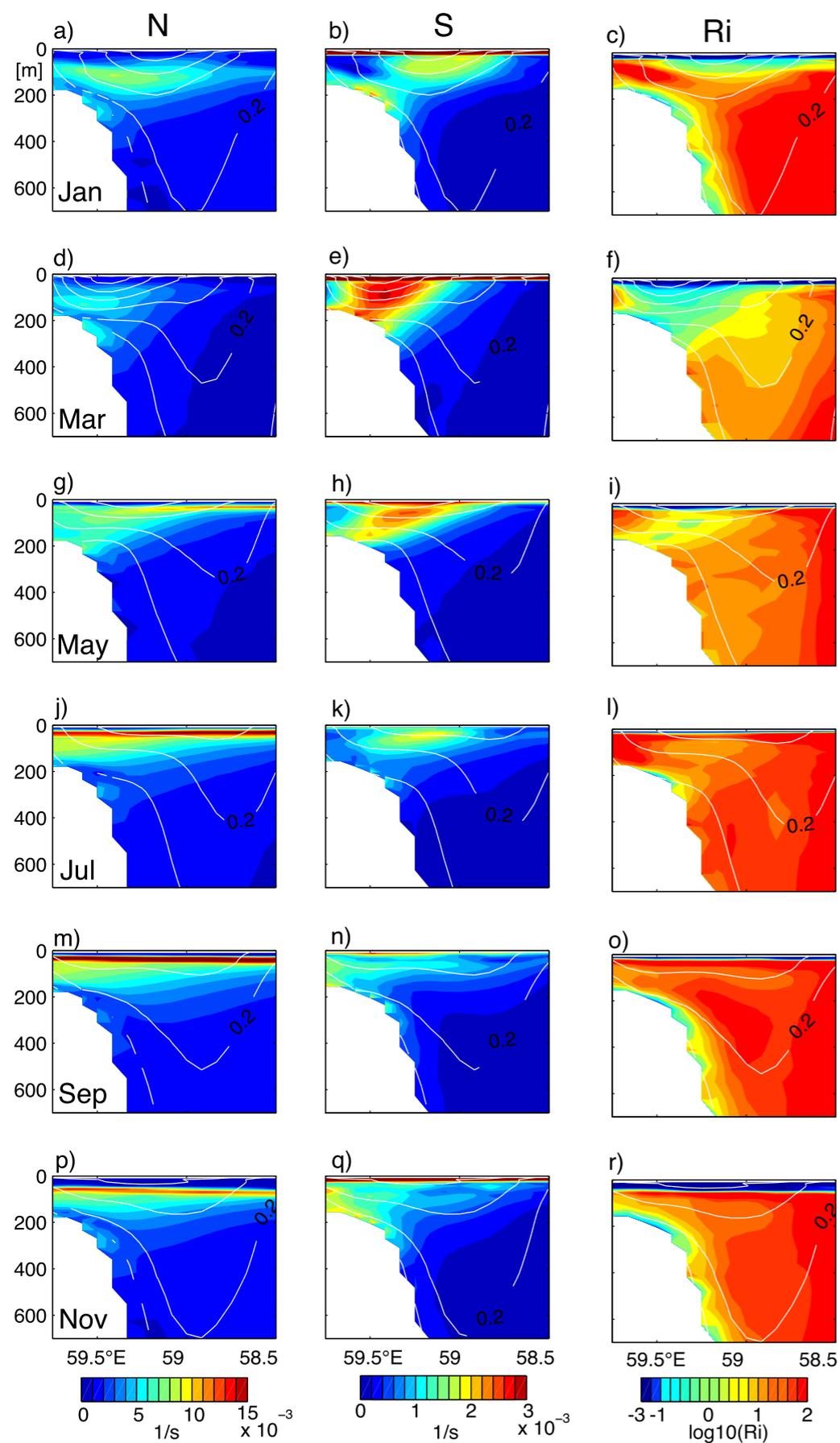


FIG. 9. Seasonal cycle of monthly-mean (left) buoyancy frequency  $N$  ( $s^{-1}$ ), (middle) vertical shear ( $s^{-1}$ ), and (right) the logarithm of  $Ri = N^2/S^2$  along  $57.6^\circ N$  from FLAME. Also shown is the alongshore velocity component [solid white lines ( $m s^{-1}$ ) with contour interval of  $0.1 m s^{-1}$ ].

# Back to the questions

1) Why does the Labrador Current become unstable during late winter and what kind of instability is at work?

-> interior mode is the **balanced mode of baroclinic instability at small Richardson numbers** (Stone 1970)

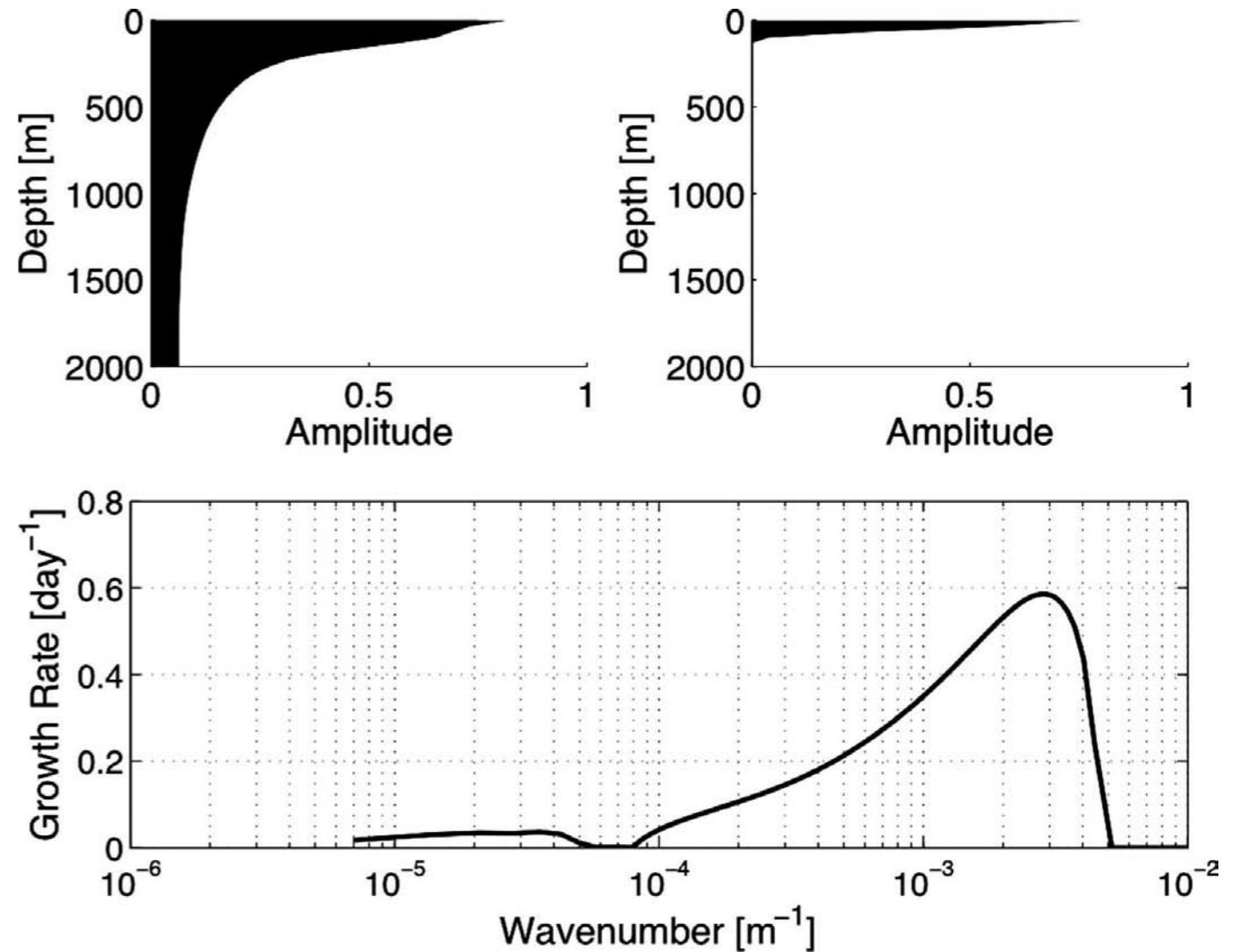
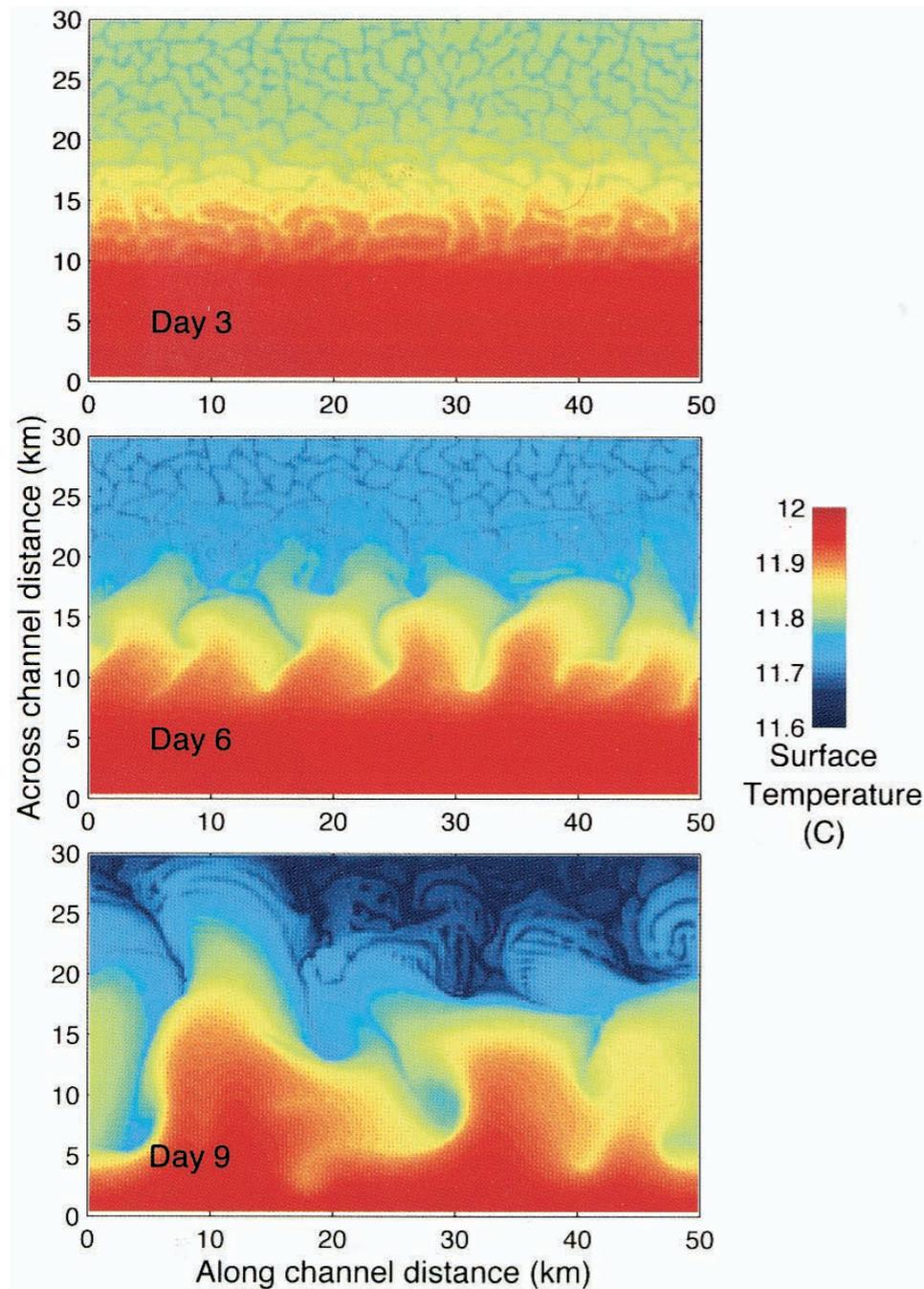
# Back to the questions

2) How do the oceanic background conditions contribute?

-> combination of **weak stratification** and **enhanced vertical shear** result in **small Richardson numbers** in late winter

-> **local EKE production** in Labrador Current by baroclinic instability at small Richardson numbers

# Baroclinic mixed layer instabilities



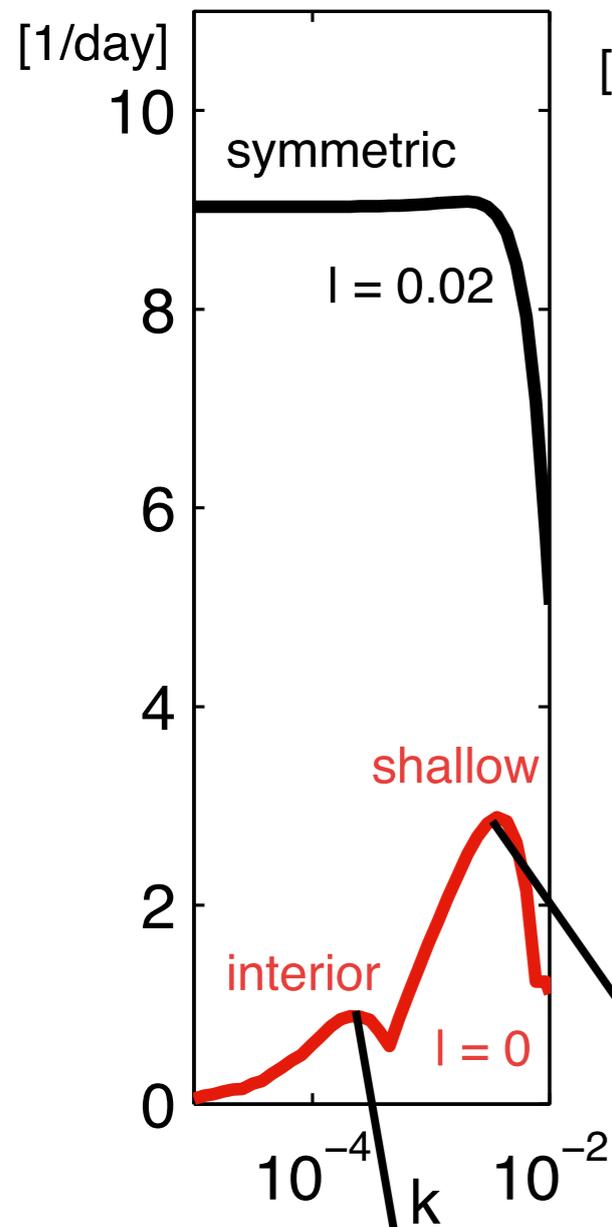
Boccaletti et al. 2007

FIG. 8. The evolution of temperature at a depth of 65 m for days 3, 6, and 9 of experiment 3. The switch over from small-scale gravitational convection to finite amplitude baroclinic instability and geostrophic turbulence is clear.

Haine and Marshall 1998

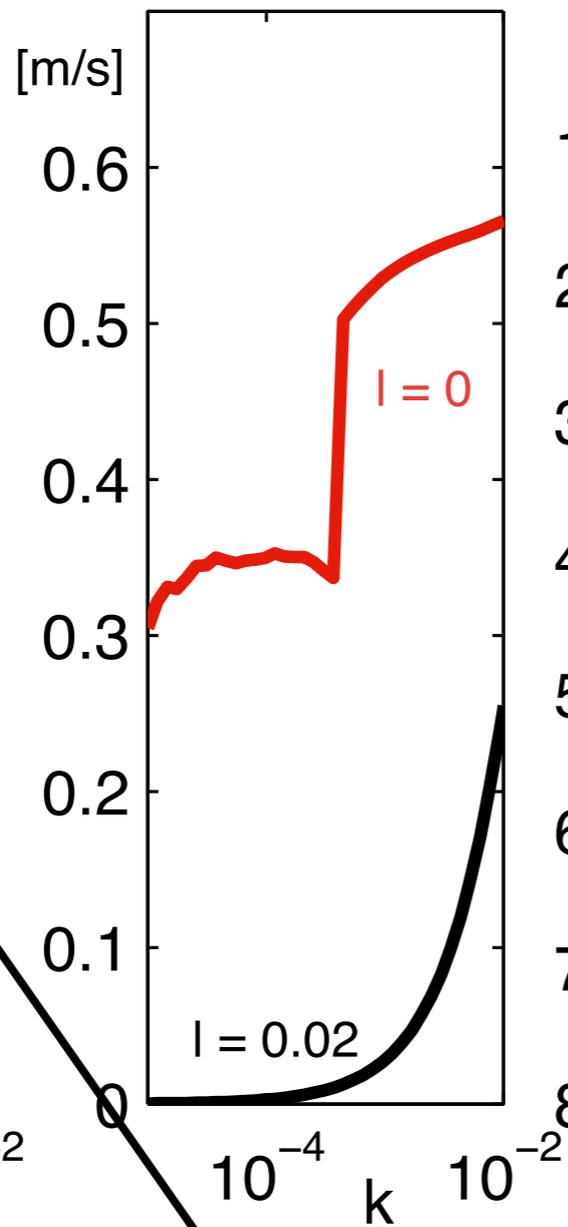
# Interior, shallow and symmetric mode

(a) growth rate



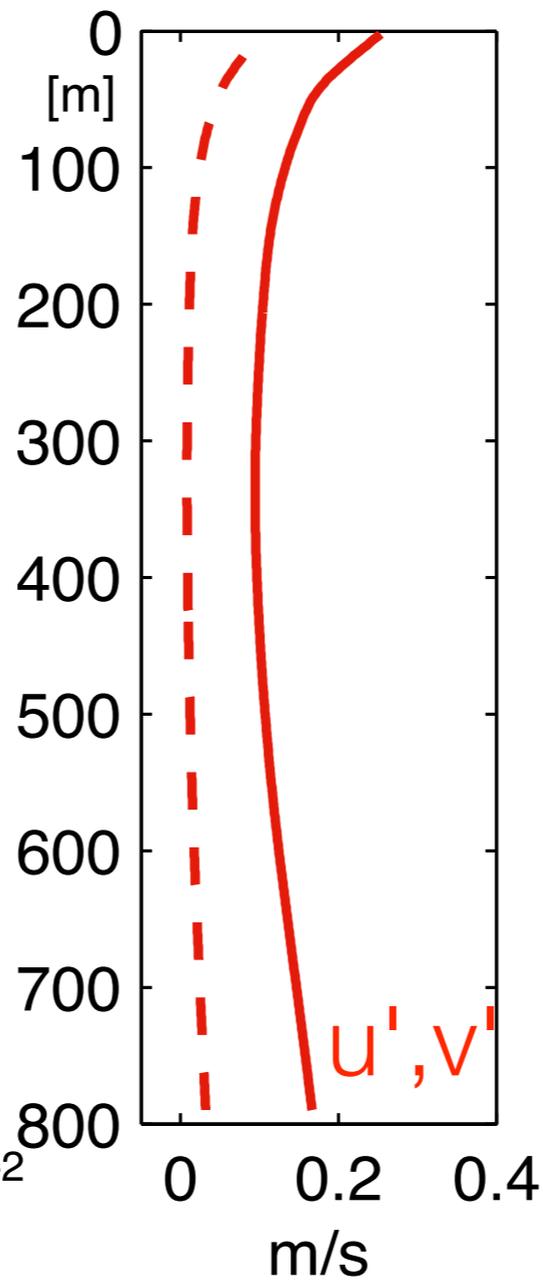
O(30-45 km)

(b) phase velocity

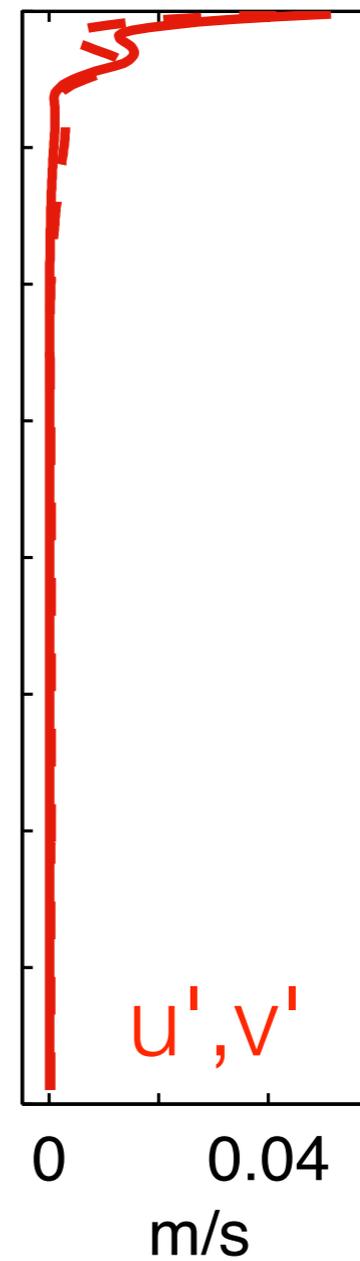


O(1 km)

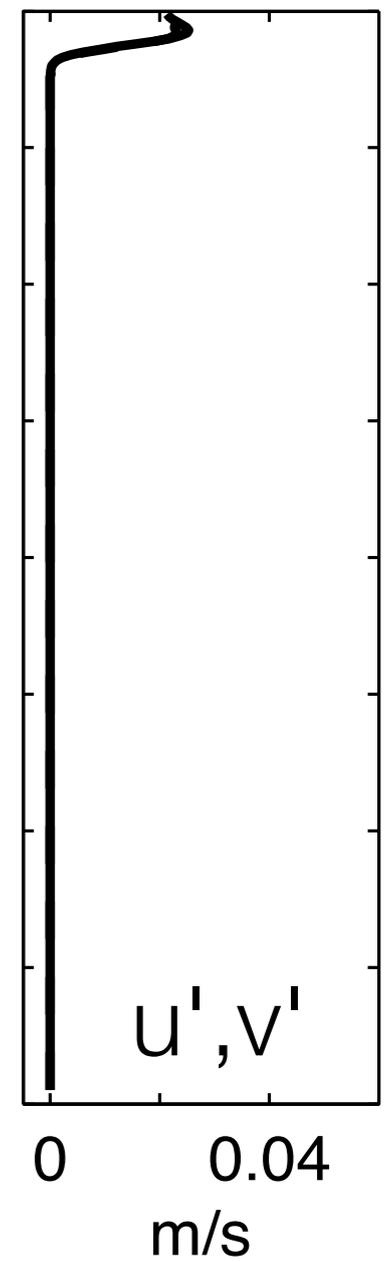
(c) interior



(d) shallow



(e) symmetric



# Summary

- interior mode  $L = O(30 - 45 \text{ km})$  has similar wavelength as first wave like disturbances in the model simulation
- weak stratification and enhanced vertical shear result in small Richardson numbers in the Labrador Current during late winter
- > growth rates of baroclinic waves increase
- > local EKE production in Labrador Current due to baroclinic instability at small Richardson numbers

# Conclusion

- **climate** and **coarse-resolution** ocean models are not able to simulate the instability process along the Labrador Current
- frontal exchange processes might be underrepresented even in high-resolution ocean models as shallow / symmetric mode claim for even higher resolution
- > or better **parametrizations** of eddy fluxes (e.g. predict eddy fluxes with linear stability analysis (e.g. Eden 2011, 2012, Vollmer et al. 2013))

# References

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Brandt et al. 2007

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doi: [http://dx.doi.org/10.1175/1520-0485\(1999\)029<2065:ODCITL>2.0.CO;2](http://dx.doi.org/10.1175/1520-0485(1999)029<2065:ODCITL>2.0.CO;2)

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Stone, P. H. (1966), On non-geostrophic baroclinic stability, *Journal of the Atmospheric Sciences*, 23(4), 390–400, doi:10.1175/1520-0469(1966)023<0390:ONGBS>2.0.CO;2.

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Stone, P. H. (1971), Baroclinic stability under non-hydrostatic conditions, *Journal of Fluid Mechanics*, 45(04), 659–671, doi:10.1017/S0022112071000260.

Straneo et al. 2002

Pickart et al. 1997, 2002,

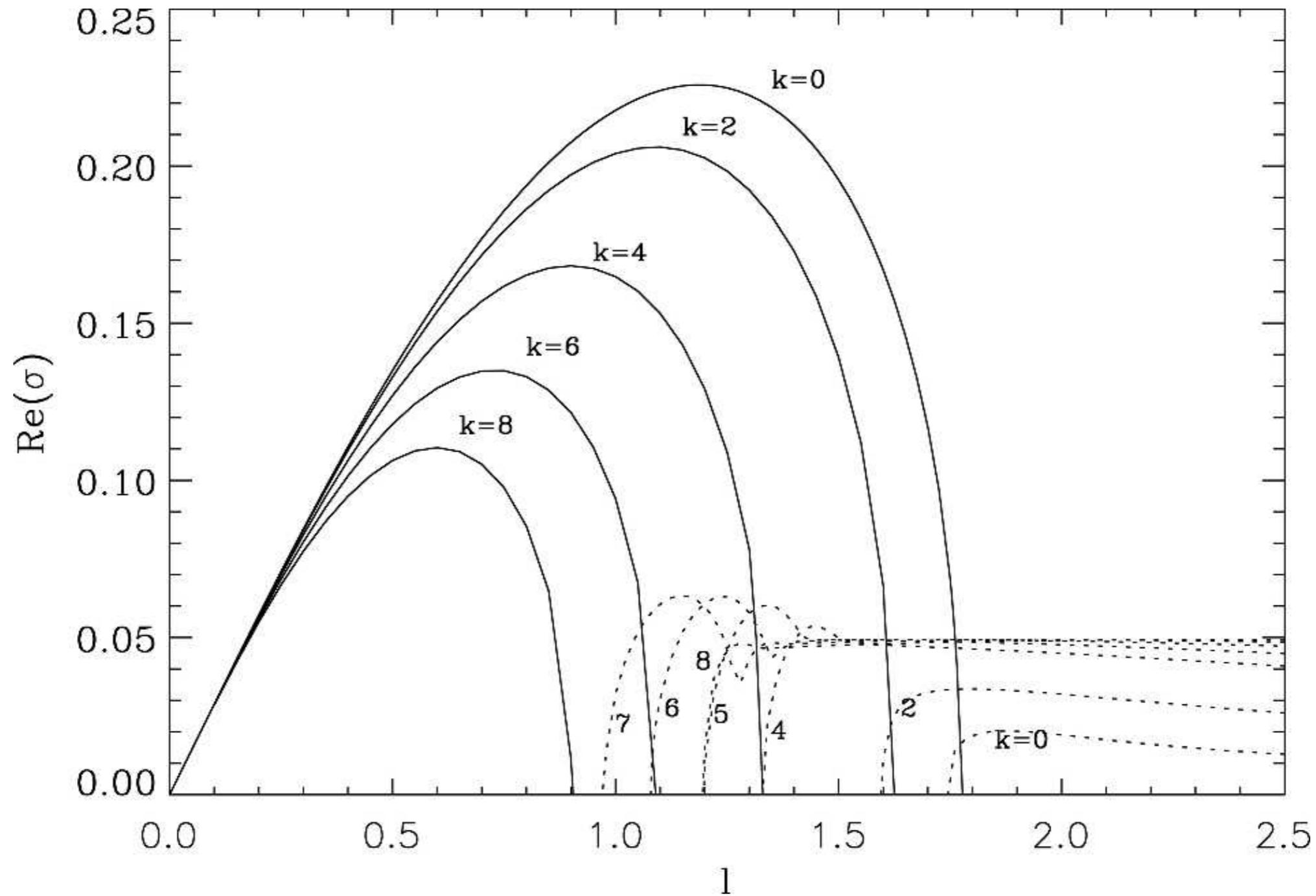
Spall and Pickart 2001

Pickart and Spall 2007

Fischer et al. 2010, 2014

Rhein et al. 2015

# Balanced vs. ageostrophic mode



Stone 1971, Molemaker 2005

# Shallow and symmetric mode

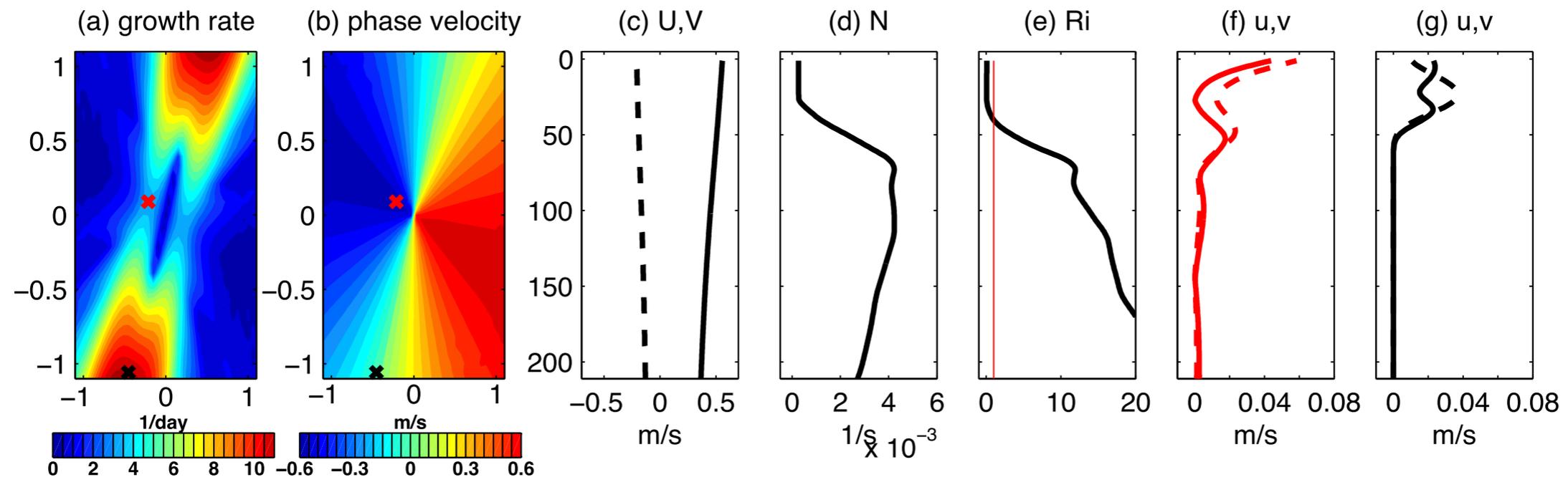


FIG. 6. The (red cross) shallow and (black cross) symmetric modes at K6 for background shear and stratification taken from January-mean values in FLAME. Shown are the (a) growth rate ( $\text{day}^{-1}$ ), (b) phase velocity ( $\text{m s}^{-1}$ ), (c) monthly-mean background velocity  $U$  (solid) and  $V$  (dashed) ( $\text{m s}^{-1}$ ), (d) background  $N$  ( $\text{s}^{-1}$ ), and (e)  $Ri$ . The red line in (e) indicates  $Ri = 1$ . (f) Velocity perturbations  $u$  (solid) and  $v$  (dashed) of the shallow mode ( $\text{m s}^{-1}$ ) and (g) the corresponding variables for the symmetric mode. The wavenumbers in (a),(b) are scaled with the mixed layer Rossby radius (see text for definition). The  $U$  and  $V$  at velocity grid points closest to the mooring positions and  $N^2$  interpolated on these points are taken as background values.

- **shallow mode** with lateral scales of  $O(1 \text{ km})$
- > submesoscale variability (Boccaletti ...)
- symmetric instabilities arise at  $Ri < 1$  (Haine ..., Straneo et al. 2002)
- both modes remain unresolved in  $1/12^\circ$  simulations